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## NAVAL POSTGRADUATE SCHOOL Monterey, California



### THESIS

ITERATIVE RETRIEVAL AND STATISTICAL SPECIFICATION OF ATMOSPHERIC THICKNESSES FROM VTPR CLEAR-COLUMN RADIANCE DATA

by

Douglas Ray Moran

Thesis Advisor:

F.L. Martin

March 1974

Thesis

Approved for public release; distribution unlimited.

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Iterative Retrieval and Statistical Specification of Atmospheric Thicknesses from VTPR Clear-Column Radiance Data

bу

Douglas Ray Moran Lieutenant, United States Navy B.S., Michigan State University, 1967

Submitted in partial fulfillment of the requirements for the degree of

MASTER OF SCIENCE IN METEOROLOGY

from the
NAVAL POSTGRADUATE SCHOOL
March 1974

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Approved by:

#### ABSTRACT

An iterative technique is developed for retrieval of thicknesses of selected atmospheric layers from VTPR "clearcolumn" radiance measurements. Layer mean temperatures for a simplified atmospheric model are retrieved by direct solution of the radiative transfer equation, and are then used to compute thicknesses of key atmospheric layers bounded by commonly used pressure levels. The retrieval technique illustrates the use of reference wave numbers that vary from layer to layer. Transmittance tuning is employed to correct systematic errors in the retrieved mean temperatures. Thicknesses of key layers retrieved by the technique from "clear-column" radiances observed during a 24 hour period at scan spots between 15 N and 45 N are separated into three latitude-band samples. Each sample is subjected to stepwise multiple regression analysis to determine the thicknessspecification of various standard layers in terms of the clear column radiances. RMS error-analyses resulting from the regression are then used to determine the quality of thickness-specifications of simulated tropospheres and stratospheres.

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#### TABLE OF SYMBOLS AND ABBREVIATIONS

	·
B <sub>i</sub> [T(p)]	Planck radiance function for temperature T at pressure level p
B <sub>i</sub> (K)	Layer mean Planck function for layer K and reference wave number $\boldsymbol{\nu}_{\boldsymbol{1}}.$
B <sub>wtd</sub> (K)	Weighted layer mean Planck value corresponding to layer reference number $\boldsymbol{\tilde{\nu}}_{K}.$
E.I.	Error index
F <sub>k</sub>	F-ratio upon entry at step k
FNWC	Fleet Numerical Weather Central
gpm	Geopotential meter
I,	Spectral radiance in channel i
mb .	millibar
NESS	National Environmental Satellite Service
NMC	National Meteorological Center
NOAA	National Oceanic and Atmospheric Administration
ν <sub>i</sub>	Wave number at center of channel i
$\tilde{v}_{K}$	Reference wave number for layer K
p <sub>o</sub>	Pressure at top of the atmosphere (.01 mb)
ps	Pressure at surface of the earth (1000 mb)
RTE	Radiative transfer equation
S.E.	Standard error of estimate
SIRS	Satellite Infra-Red Spectrometer
SR	Scanning Radiometer
SST	Sea surface temperature
σ	Standard deviation
T(K)	Mean temperature of layer K

τ <sub>1</sub> (p)	Fractional transmittance of atmosphere in channel i from level p to $p_0$						
VTPR	Vertical Temperature Profile Radiometer						
R	Multiple correlation coefficient						

#### ACKNOWLEDGEMENTS

The author wishes to express his appreciation to Professor Frank L. Martin for his generous assistance and guidance in the research and preparation of this paper.

Appreciation is also expressed to the author's family for patience and understanding during the period of thesis work, and to the staff of the W. R. Church Computer Facility for assistance in computer phases of the research.

#### I. INTRODUCTION

Retrieval of meteorological parameters from satellite radiance measurements has been the object of numerous research studies since Kaplan [1959] demonstrated that vertical temperature profiles of the atmosphere could be inferred from satellite spectral radiance observations in the 15 µm band of carbon dioxide emission. Application of retrieval techniques to satellite measurements has been possible since the launch of NIMBUS III in 1969 and to date emphasis has been on retrieval of temperature profiles.

However, for purposes of numerical weather prediction, the atmospheric thickness of a specified pressure interval is normally a more basic parameter than temperature [Fleming, 1972]. Furthermore, thicknesses of specific atmospheric pressure intervals, or layers, may be retrieved by direct solution of the radiative transfer equation, thereby eliminating the need for "a priori" statistical information necessary for retrieval by regression or inverse matrix methods [Fritz, Wark, et al., 1972].

A direct retrieval method for obtaining specific layer mean temperatures, and hence thicknesses, from satellite radiance measurements was presented by Fleming [1972]. The method proved to be too time-consuming to be operationally adaptable by the National Environmental Satellite Service (NESS) to the National Meteorological Center (NMC) analysis

scheme as described by McMillin, Wark, et al. [1973], and the retrieved thickness values showed little improvement in accuracy when compared to results obtained by a statistical regression technique which gave a T(p) profile at mandatory levels from which thicknesses were computed [Smith and Fleming, 1972]. However, the concept proposed by Fleming [1972] appeared to have at least diagnostic merit for application to vertical analysis schemes such as that of Fleet Numerical Weather Central (FNWC), which employs a layer thickness as an input parameter [Holl, et al., 1964].

Fleming proposed a method whereby the unit square wave function denoted by  $R(\ln p/p_s)$  could be determined as a linear combination of spectral radiances,  $I_i$ , with coefficients,  $c_i$ , chosen to minimize the right side of

$$\sum_{i=1}^{N} c_{i}I_{i} = \Delta \overline{T} \int_{p_{o}}^{p_{s}} R[\ln p/p_{s}; p_{1}, p_{2}]d(\ln p) . \qquad (1)$$

Here

 $\Delta \overline{T}$  = mean temperature for pressure interval (p<sub>1</sub>,p<sub>2</sub>).

p<sub>o</sub>, p<sub>s</sub> = pressures at top of atmosphere and surface of earth.

The R-functions were sought to provide, in sequence, exact fits to the square wave in the significant layers of the atmosphere. By Fleming's theory they should ideally fit the square wave of Fig. 1 between  $p_1$  and  $p_2$  and be zero elsewhere along the p-axis.

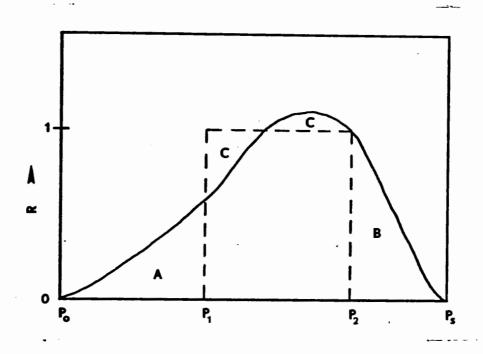


FIG. 1. Exact vs. calculated rectangular function R. Desired solution indicated by dashed line, calculated solution indicated by solid line.

In practice, the calculated R functions did not fit the square wave exactly and left errors in adjacent pressure regions as depicted by areas A and B of Fig. 1. Inexactness of fit at the top of the square wave resulting in area C was considered unimportant because it tended to average out to the required thickness constraint over  $(p_1,p_2)$ .

The implication was that the better the R function fits the square wave for a given layer  $(p_1,p_2)$ , the better the mean temperature of the layer was specified by the radiance measurements.

The object of this thesis was first to develop a direct retrieval technique for layer-mean temperatures (and therefore thicknesses) which required a minimum of computer time, and could conceivably be less sensitive to channel noise than Fleming's method. A second objective was designed to test retrieved thicknesses for optimum layer specification by statistical methods, as implied by Fleming's square wave concept.

In this thesis a new simplified iterative retrieval technique was developed for use with VTPR "clear-column" radiances to obtain mean temperatures over greater pressure intervals than the mandatory pressure increments used by Fleming. An iterative retrieval technique was used because of the minimum computer time and space required. The use of larger pressure intervals was employed to reduce the effect of channel noise and to improve accuracy, since the accuracy of mean temperatures derived from radiance measurements usually improves with increasing pressure interval [Hayden, 1971].

Key-layer thicknesses derived hypsometrically from the mean T(p) profile were then subjected to stepwise multiple regression analysis to determine which key layer was best specified by the "clear-column" radiances. The troposphere and lower stratosphere were considered separately, that is layers crossing the tropopause level such as 500 to 50 mb were not considered.

Results of the regression analysis were examined first to determine which particular layer in the troposphere (and also which layer in the lower stratosphere) could be used as the most effective thickness from VTPR retrieval for use in the FNWC vertical-structure analysis scheme, which currently makes use of the 1000 to 300 mb thickness in processing of conventional sounding data.

Next, combinations of sub-layers which span the troposphere were examined to determine if a better statistical fit for the retrieved tropospheric thickness could be obtained from the combinations of sub-layers, and if so, then which combination was best. A similar study was made for the lower stratosphere.

#### II. SATELLITE DATA

The launch of NOAA-II with its Vertical Temperature

Profile Radiometer (VTPR) instrument in October, 1972

marked a major improvement in radiance measurement capability.

The VTPR instrument is superior to its earlier counterparts,
the SIRS-A of NIMBUS III and the SIRS-B of NIMBUS IV, in that
it has much better spatial resolution. Sub-satellite dimensions for scans spots are approximately 69 by 67 km for VTPR
compared to 225 by 225 km for SIRS-A and SIRS-B. The
improved resolution permits more accurate computation of
"clear-column" radiances, which are equivalent to radiances
that would be observed in completely clear skies. Use of
"clear-column" radiances for retrieval eliminates the necessity to correct for cloud cover which was a significant
problem in retrieval techniques developed for use with SIRS-A
and SIRS-B data.

NOAA-II orbits the earth every 115 minutes at an altitude of 1464 km. Figure 2 illustrates the earth projection of seven orbits. North to south portions of the orbit are indicated by solid lines; south to north portions are indicated by dashed lines. Shaded areas depict areal coverage during two orbits. North to south equator crossings occur at 0900 and south to north crossings occur at 2100 local solar time.

The VTPR instrument scans perpendicular to the satellite path in 23 discrete steps from left to right, representing

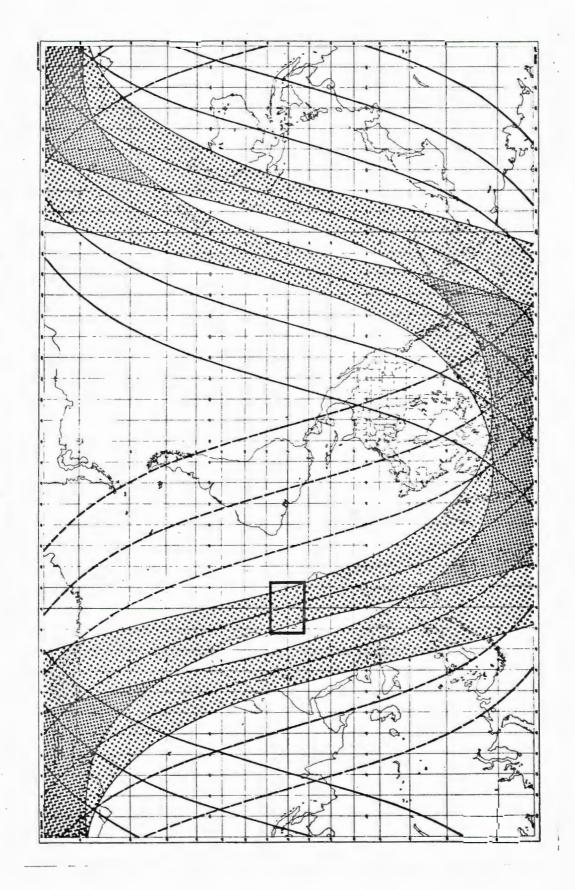


FIG. 2. Satellite tracks for NOAA II VTPR coverage

a scan path of 30.3 degrees both sides of the local nadir. Each step provides a "scan spot" which is observed by the VTPR instrument in six spectral intervals of the 15 µm band of carbon dioxide, in one interval of the 12 µm atmospheric window, and in one interval of the 19 µm water vapor band.

The eight spectral intervals, or channels, along with their respective half-widths and central wave numbers are listed in Table 1.

TABLE 1. VTPR channel designators corresponding half-widths, and central wave numbers.

Channel	1	2	3	4	5	6	7	8
Wave No.(v <sub>i</sub> ) (cm-1)	668.5	677.5	695.0	708.0	725.0	747.0	535.0	835.0
Half-width (cm-1)	3.5	10.0	10.0	10.0	10.0	10.0	18.0	10.0

Channels one through six are carbon dioxide channels, channel seven is the water vapor channel, and channel eight is the window channel. Maximum relative measurement error between any two channels except the  $668.5 \text{ cm}^{-1}$  channel is  $0.25 \text{ mW/(m}^2 \text{ ster cm}^{-1})$ ; maximum relative error between the  $668.5 \text{ cm}^{-1}$  channel and any other channel is  $0.75 \text{ mW/(m}^2 \text{ ster cm}^{-1})$ .

Other instruments aboard NOAA-II include a two-channel Scanning Radiometer (SR) which measures radiances in the 10.4-12.5 and 0.5-0.7 µm intervals. Resolution is more

refined in the SR scan spots than that of the VTPR, with subsatellite dimensions at the nadir being approximately 7.5 by 7.5 km. Statistical techniques are used to identify scanning radiometer measurements which signify cloud-free areas, and these "clear-column" SR window channel radiances are then used to determine sea surface temperatures for the VTPR scan spots.

Conversion of the VTPR raw radiance measurements to "clear-column" radiances is accomplished by first dividing scan spots into analysis arrays. Scan spots from eight successive scan lines are divided into three boxes, or subarrays, of 8 by 8, 8 by 7, and 8 by 8 spots as illustrated in Fig. 3, which is an enlargement of the boxed area outlined in Fig. 2. From the VTPR raw radiance measurements and the SR derived sea surface temperatures of scan spots within each sub-array, a single set of "clear-column" radiances is computed by statistical methods [McMillin, Wark, et al, 1973], and assigned to central scan spot locations indicated by X's in Fig. 3.

The procedure is essentially to compute an 835.0 cm<sup>-1</sup> window channel radiance by the Planck formula using the sea surface temperature, and to compare the computed value against the measured 835.0 cm<sup>-1</sup> window channel radiance. If the measured 835.0 cm<sup>-1</sup> radiance value equals or exceeds the computed value (from the known sea-surface temperature field, denoted by SST), the radiances are considered to be in agreement and the scan spot is assumed to be cloud-free.

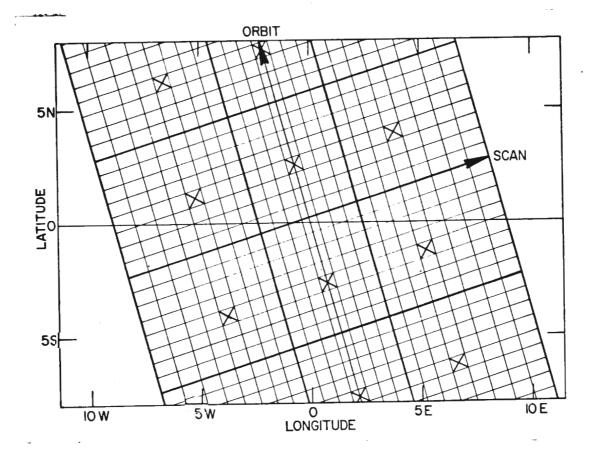


FIG. 3. VTPR scan pattern and data analysis array for the outlined box in Fig. 2.

If the computed value (from the SST) exceeds the measured value, the scan spot is assumed to contain significant cloud cover and equivalent "clear-column" radiance values are calculated by noting that for a given wave number,  $\nu_{\bf i}$ ,

$$\frac{I_{clr}(v_i) - I_2(v_i)}{I_{clr}(v_8) - I_2(v_8)} = \frac{I_{clr}(v_i) - I_1(v_i)}{I_{clr}(v_8) - I_1(v_8)}.$$
 (2)

Here

 $I_{clr}(v_8)$  = window radiance computed from the SR derived sea surface temperature.

 $I_1(v_1), I_2(v_1) = \text{raw radiances in channel i measured at scan spots 1 and 2 having the same seasurface temperature.}$ 

 $I_{clr}(v_i) = desired "clear-column" radiance value for wave number .$ 

 $I_1(v_8), I_2(v_8)$  = radiance measured for the 835.0 cm<sup>-1</sup> window channel at scan spots 1 and 2.

As shown in Fig. 4, the three points  $[I_{clr}(v_i), I_{clr}(v_8)]$ ,  $[I_1(v_i), I_1(v_8)]$ , and  $[I_2(v_i), I_2(v_8)]$  lie on the straight line determined from the measured values of  $I_1(v_i)$ ,  $I_1(v_8)$ ,  $I_2(v_i)$ , and  $I_2(v_8)$ . The slope of this line is expressed by the right side of eq. (2).

The value of  $I_{\rm clr}(\nu_i)$  can then be determined from the known value of  $I_{\rm clr}(\nu_8)$ . When computing  $I_{\rm clr}(\nu_i)$  values, radiances from adjacent scan spots with different nadir angles are adjusted to a common zenith angle.

After clear radiance values are obtained for all channels at each scan spot within a sub-array, all values from adjacent scan spots are examined. The sub-array maximum radiance set yields a single set of eight "clear-column" radiances, and a sea surface temperature positioned at the center of the sub-array. This results in reduction of raw radiances for the 184 scan spots of the 8 by 23 analysis array, to "clear-column" radiances at only the centers of the three sub-arrays.

The resulting VTPR "clear-column" radiance values and associated data are recorded on archival tapes. An archival tape for April 12, 1973, was provided for this study through

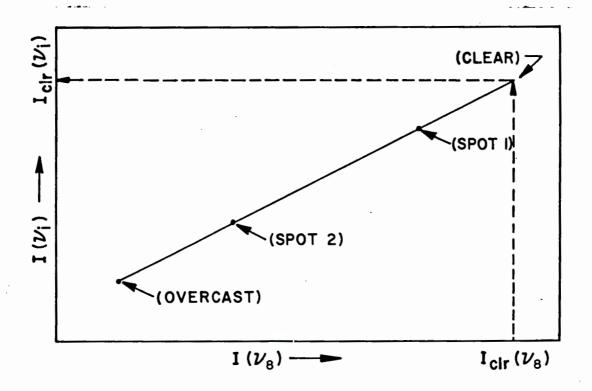


FIG. 4. Procedure for determining "clear-column" radiances.

the kind auspices of Dr. D. Q. Wark of NESS. From a printout of archival file II (the "clear column" file), radiances for the six carbon dioxide channels and the sea surface temperature, as well as the geographic coordinates, were extracted for clear column scan spots between 15 and 45 degrees north latitude.

Data from the tape printout were then transformed to correct dimensions as follows:

```
N. Latitude in degrees = (tape\ value)\ x\ .10\ -90
W. Longitude in degrees = (tape\ value)\ x\ .10
Radiances in mW/(m<sup>2</sup> ster cm<sup>-1</sup>) = (tape\ value)\ x\ .05
Sea surface temperature in °K = (tape\ value)\ x\ .20\ +\ 269.9
```

Additional information on characteristics of the NOAA-II satellite and its VTPR and SR radiance data may be found in NOAA technical reports [Fritz, Wark, et al., 1972] and [McMillin, Wark, et al., 1973] from which most of the details of this section were taken.

#### III. RETRIEVAL TECHNIQUE

Given VTPR "clear-column" radiance values for channels one through six and the sea surface temperature computed essentially from SR measurements, corresponding layer mean temperature profiles were obtained using an iterative technique derived from that initially presented by Smith [1970] and modified by Martin [1973] for the purposes of this study. Mean temperatures are then converted to layer thicknesses by use of the hypsometric equation, as if the mean temperatures of the iterative procedure applied to their central pressure levels.

#### A. MATHEMATICAL DEVELOPMENT

For a cloudless, non-scattering atmosphere in thermo-dynamic equilibrium, the spectral radiance observed at the top of the atmosphere for each channel is related to the vertical temperature profile and absorbing gas structure by the <u>radiative transfer equation</u> (RTE) [Fritz, Wark, et al., 1972]:

$$I_{i} = B_{i}[T(p_{s})]\tau_{i}(p_{s}) + \int_{x(p_{s})}^{x(p_{o})} B_{i}[T(p)] \frac{d\tau_{i}(p)}{dx(p)} dx(p)$$
(A) (B)

where

 $I_1 =$  spectral radiance in channel i, (i=1,2,...,6),

- $\tau_1(p)$  = fractional transmittance of the atmospheric  $co_2$  in channel i from pressure level p to  $count_2 = .01 \text{ mb}$ ,
  - x(p) = an arbitrary function of pressure which behaves in the vertical similar to log  $p/p_0$ .

Term (A) is the atmospheric transmittance of the Planckian radiance from the surface of the earth. Term (B) is the atmospheric contribution to the radiance. Subscripts s and o refer to surface of the earth (1000 mb) and top of the atmosphere (.01 mb), respectively.

When 100 pressure levels are linearly scaled by  $p^{2/7}$  and the result adopted for x(p), it follows that

$$p(J) = .01[1 + (J-1)(0.26087836)]^{7/2}$$
 (4)

and J = 1, 2, ..., 100 are pressure levels numbered from top of the atmosphere to surface of the earth, the radiative transfer equation can be rewritten [Martin, 1973]

$$I_{1} = B_{1}[T(100)]_{\tau_{1}}(100) + \int_{J=100}^{J=01} B_{1}[T(J)] \frac{d\tau_{1}(J)}{dJ} dJ$$
 (5)

The Planck radiance function is defined as

$$B_{1}[T(J)] = C_{1}v_{1}^{3}/[e^{(C_{2}v_{1}/T(J))} - 1]$$
 (6)

where

 $v_i$  = wave number for channel i,  $c_1$  = 1.9061 x 10<sup>-5</sup> erg cm<sup>2</sup> sec<sup>-1</sup> ster<sup>-1</sup>,  $c_2$  = 1.43868 cm °K.

When the 100 pressure levels are combined into the 17 atmospheric layers depicted in Fig. 5, eq. (5) can be evaluated in quadrature form

$$I_{i} = B_{i}[T(100)]\tau_{i}(100) + \sum_{K=1}^{17} \overline{B_{i}}(K)\Delta\tau_{i}(K)$$
 (7)

where

$$\overline{B_{i}}(1) = \frac{1}{6} \{B_{i}[T(01)] + 4B_{i}[T(02.5)] + B_{i}[T(04)]\}$$
 (8)

is the layer mean Planck function for layer K = 1, and

$$\overline{B_{1}}(K) = \frac{1}{6} \{B_{1}[T(J_{K}-3)] + 4B_{1}[T(J_{K}) + B_{1}[T(J_{K}+3)]\}$$
 (9)

is the layer-mean Planck function for layer K > 1. In (7),  $\Delta \tau_{\bf i}(K) \text{ is the transmittance in channel i from layer K as}$  defined by

$$\Delta \tau_{i}(K) = \tau_{i}(J_{K}-3) - \tau_{i}(J_{K}+3)$$
. (10)

Finally,  $J_k$  is the value of J at the center of layer K.

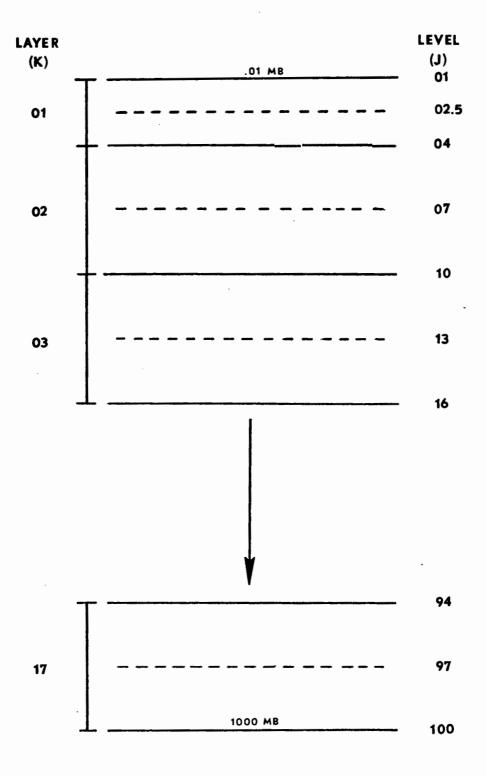


FIG. 5. Seventeen K-layer atmosphere model used for mean-temperature retrieval. Note that the top layer spans only four J-levels while all others span seven J-levels.

Equation (7) can be written in iterative form as

$$I_{i} - I_{i}^{n} = \{B_{i}^{n+1}[T(100)] - B_{i}^{n}[T(100)]\}_{\tau_{i}}(100)$$

$$+ \sum_{K=1}^{17} \{\overline{B}_{i}^{n+1}(K) - \overline{B}_{i}^{n}(K)\}_{\Delta \tau_{i}}(K)$$
(11)

where

I<sub>i</sub> = observed radiance for channel i, considered to
 be the final iterative value,

 $I_i^n$  = calculated radiance for channel i from eq. (7) at iteration number n.

Then, following Smith [1970], in each channel the difference

$$\{\overline{B}_{\mathbf{i}}^{n+1}(K) - \overline{B}_{\mathbf{i}}^{n}(K)\}$$
 (12)

is independent of pressure within all atmospheric layers; hence the following iterative equation is obtained from eq. (11):

$$\overline{B}_{i}^{n+1}(K) = \overline{B}_{i}^{n}(K) + [I_{i} - I_{i}^{n}] . \qquad (13)$$

From a first guess temperature profile T(J), J=1, 2, 2.5, 3,..., 100, Planck radiance values,  $B_{\bf i}[T(J)]$  for all channels at each level can be computed from eq. (6). These values can then be used to determine layer mean Planck (1) values,  $\overline{B}_{\bf i}$  (K), by eqs. (8) and (9), and subsequently radiances,  $I_{\bf i}$ , in accordance with eq. (7). The difference

between observed and calculated radiances,  $[I_i - I_i^{(1)}]$ , can be applied as a residual correction to obtain adjusted layer mean Planck values,  $\overline{B}_i^{(2)}(K)$ , by use of eq. (13), and then improved radiances,  $I_i^{(2)}$ , can be calculated from eq. (7).

By continuing the process of adjusting layer mean Planck values and calculating new radiance values until the difference between observed and calculated radiance for each channel satisfies a convergence criterion at the final iteration step N, final layer mean Planck values,  $\overline{B}_{\mathbf{i}}^{N}(K)$ , are obtained for all channels,  $i=1,\ldots,6$  and all layers  $K=1,2,\ldots,17$ .

From the final layer mean Planck values a single mean temperature for each layer can be computed by noting that the fraction of calculated radiance for channel i and layer K,  $\Delta I_{i}(K)$ , can be expressed

$$\Delta I_{\uparrow}^{N}(K) = \overline{B}_{\uparrow}^{N}(K) \Delta \tau_{\uparrow}(K) . \qquad (14)$$

A similar expression applies to each of the six channels for a given layer, and a weighted layer mean Planck value can be formed by summing the  $\Delta I_1^N(K)$  over all  $i=1,\ldots,6$ . Therefore, for a layer K the weighted layer mean Planck value,  $\overline{B}_{w+d}^N(K)$ , corresponding to a reference wave number,  $\tilde{\nu}_K$ , for the layer is computed [Martin, 1973]:

$$\overline{B}_{w+d}^{N}(K) = \sum_{i=1}^{6} \overline{B}_{i}^{N}(K) \Delta \tau_{i}(K) / \sum_{i=1}^{6} \Delta \tau_{i}(K).$$
 (15)

Substituting the weighted layer mean Planck value into the Planck equation yields

$$\overline{B}_{w+d}^{N}(K) = c_1 \tilde{v}_K^{3} / [e^{(c_2 \tilde{v}_K / T(K))} - 1]$$
 (16)

where  $\tilde{\nu}_K$  is the reference wave number and T(K) is the mean temperature for layer K at the first guess step of iteration.

The reference wave number for a layer K can be determined by evaluating the weighted layer mean Planck value from eq. (16) using the initial layer mean Planck values  $B_{\bf i}^{(1)}(K)$ , substituting the value of  $\overline{B}_{\bf wtd}^{(1)}(K)$  thus obtained into a rearranged form of (16)

$$\overline{B}_{w+d}^{(1)}(K) \left[ e^{(C_2 \tilde{v}_K / T(K))} - 1 \right] - C_1 \tilde{v}_K^3 = 0 , \qquad (17)$$

and solving for  $\tilde{\nu}_K$  using the Bailey iteration method of solution of transcendental differential equations [McCalla, 1967]. Repetition for each layer gives reference wave numbers for all layers,  $\tilde{\nu}_K$ ,  $K = 1, 2, \ldots, 17$ .

The layer mean temperatures for each layer can then be calculated from the weighted layer mean Planck value corresponding to the final layer mean Planck value,  $\overline{B}_{\mathbf{i}}^{N}(K)$ , by

$$\overline{T}(K) = C_2 \tilde{v}_K / \ln \left[ \frac{C_1 \tilde{v}_K^3 + \overline{B}_{w+d}^N(K)}{\overline{B}_{w+d}^N(K)} \right]$$
 (18)

where  $\tilde{\nu}_K^{}$  is the reference wave number computed only at the initial step in accordance with (17).

Layer-mean temperatures may then be converted to layer-thickness values by integrating the hydrostatic equation between top and bottom pressure levels of the layer to obtain the hypsometric equation [Haltiner and Martin, 1957]

$$\Delta Z = \frac{R_d}{g} \overline{T}(K) \ln(\frac{p_1}{p_2})$$
 (19)

where

 $\Delta Z$  = thickness of a layer in meters,

 $R_d = 0.287 \text{ joules/(gm °K)},$ 

 $g = 9.80 \text{ m/sec}^2$ ,

 $\overline{T}(K)$  = mean temperature of the layer in  ${}^{\circ}K$ ,

 $p_1$ ,  $p_2$  = pressures at top and bottom of layer in mb.

#### B. APPLICATION

#### 1. Retrieval Input Parameters

In addition to the satellite radiance data already discussed in Section III, first guess temperature profiles for each scan spot and atmospheric transmittance values for all six channels are required as input parameters for the retrieval program.

#### a. First Guess Temperature Profiles

A first guess temperature profile for each scan spot was derived from 56-level climatological profiles drawn from the U.S. Standard Atmosphere Supplement [1966]. The 15 N annual profile was assumed to be representative of an April profile at that latitude. Profiles for 30 N and

45 N for both January and July were interpolated with respect to time to give equivalent April profiles at each latitude. The three resulting "April" climatological profiles were then expanded to 100 level profiles by interpolating temperature with respect to pressure to give temperatures at the  $p^{2/7}$  or J-levels defined by eq. (4).

The 56-level climatological pressure levels and corresponding temperatures for the 15 N as well as the 30 N and 45 N January and July climatological profiles are included in Appendix A, along with the interpolation scheme used for expansion from 56 to 100 level profiles.

Given the latitude of a scan spot, the corresponding first guess profile was obtained by interpolating with respect to latitude only between the 100 level "April" climatological profiles north and south of the scan spot, and by interpolating with respect to pressure between J=2 and J=3 to obtain the J=2.5 level temperature. The 1000 mb climatological temperature was then replaced by the sea-surface temperature to "tie down" the profile at the lower boundary J=100.

#### b. Atmospheric Transmittance Values

The absorbing gas structure for the six carbon dioxide channels was assumed to be represented by the 100 J-level transmittances listed in Appendix B. These are transmittances calculated for a model atmosphere and the standard temperature profile which is also included in the

Appendix. However, as will be discussed in Section V, some of the transmittance profiles were later adjusted to improve retrieval results.

#### 2. Computational Procedure

In practice, the reduction of VTPR "clear-column" radiances to atmospheric thicknesses was accomplished using two separate computer programs: one to retrieve mean temperatures for all scan spots considered, and a second to convert mean temperatures to thicknesses as well as to sort data by latitude band for analysis by the BIMED 02R regression program. The two programs just described appear with sample outputs immediately following Appendix B.

#### a. Mean Temperature Retrieval

The procedure for retrieving layer mean temperatures for layers K = 1,2,...,17 from the VTPR "clear column" carbon dioxide radiances can be summarized as follows:

- [1] Derive a first guess temperature profile T(J), J = 1,2,2.5,3,...,100 in the manner described previously, and compute Planck radiance values,  $B_{i}[T(J)]$ , for each J-level using eq. (6).
- [2] Calculate layer mean Planck values,  $\overline{B}_1^{(1)}(K)$ ,  $K=1,2,\ldots,17$  in accordance with eqs. (8) and (9).
- [3] Compute reference wave numbers for each layer using the initial layer-mean Planck values to form the weighted layer mean Planck value,  $\overline{B}_{w+d}^{(1)}(K)$  by use of eq. (15) and then solving (17) for  $\tilde{\nu}_K$  by the Bailey iterative method.

- [4] Use the layer-mean Planck values to calculate the nth iterative radiances,  $I_1^n$ , i = 1, 2, ..., 6 by eq. (7).
- [5] Compare the observed radiances with the calculated radiances and apply the difference,  $[I_i I_i^n]$ , as a residual correction to adjust the layer-mean Planck values in accordance with (13).
- [6] Repeat steps [4] and [5] until convergence is achieved, convergence being defined as that condition in which

$$\left|\frac{\left[I_{\underline{1}} - I_{\underline{1}}^{N}\right]}{I_{\underline{1}}}\right| < 0.0001 \tag{20}$$

- [7] From the final layer-mean Planck values,  $\overline{B}_{\bf i}^N(K)$ , compute the corresponding layer-mean temperatures using eq. (18).
  - b. Standard-layer Thickness Calculation

Instead of calculating thicknesses for the 17 layers for which mean temperatures were retrieved, it was decided to compute thicknesses between standard pressure levels that would permit a more complete and orderly examination of layer combinations for specification of the troposphere and stratosphere as will be discussed in Section VII. The standard pressure levels chosen were the 56-climatological profile levels plus the additional four levels at 650, 550, 450, and 80 mb.

Computation of thicknesses for the larger number of pressure intervals was accomplished by the procedure summarized as follows:

- [1] Consider the 17-layer mean temperatures obtained by retrieval to be located at the mid-levels of each layer, that is at levels  $J = 2.5, 7, 13, \ldots, 97$ .
- [2] "Tie down" the resulting profile to the sea surface temperature at level J = 100 (1000 mb).
- [3] Determine the lapse rate between levels J = 7 and J = 2.5 and continue the lapse rate to level J = 01 (.01 mb) to determine the temperature at the top of the atmosphere. This completes a 19-level profile with temperatures at J-levels 1, 2.5, 7, 13, ..., 97, 100.
- [4] Interpolate for temperature with respect to pressure, to derive temperatures at the 60 standard pressure levels from the 19-level profile on the J-scale.
- [5] Calculate thicknesses of layers between successive standard pressure levels using a modification of the hypsometric equation (19)

$$\Delta Z = \frac{R_d}{g} \frac{(T_1 + T_2)}{2} \quad ln \ (\frac{p_1}{p_2})$$
 (21)

where

 $T_1$ ,  $T_2$  = temperatures (°K) at top and bottom pressure levels of the layer under consideration.

## IV. EVALUATION OF RETRIEVAL ACCURACY

If matching radiosonde soundings had been available, the retrieved mean temperatures for a given "clear-column" scan spot could have been compared to mean temperatures calculated from the corresponding radiosonde profile. However, matched radiosonde data were not available.

Therefore to evaluate accuracy of the retrieved mean temperatures, profiles constructed from the retrieved temperatures were first compared with the climatologically derived first-guess profiles to determine whether or not there were systematic differences between the two. Next, lapse rates between the bottom and top of layers of thickness  $\Delta J = 3$  were examined and layers having super-adiabatic lapse rates were identified.

Retrieved profiles were constructed by assigning mean temperatures to pressure levels at the middle of each layer by extending the lapse rate between levels J=7 and J=2.5 to level J=1 to compute a temperature for the top of the atmosphere, and by "tying down" the profiles by setting the temperature at J=100 equal to the sea surface temperature. With temperatures then fixed for the 19 levels J=1, 2.5, 7, 13, ..., 97, 100, the temperatures at convenient intermediate levels were computed by interpolation to give the profiles T(J), J=1, 2.5, 4, 7, ..., 94, 97, 100, defined now at 34 J-levels for each scan spot.

Initial retrieval attempts resulted in profiles that were systematically colder than first guess profiles by a few degrees at all levels except J = 100 where the retrieval temperature was set equal to the sea surface temperature. The important bias upon retrieval was that super-adiabatic lapse rates of from 12 to 18 °K per km were observed between 1000 and 900 mbs in virtually every retrieved temperature profile. Lapse rates above 900 mbs were approximately the same as those of the profiles derived from climatological standards.

Similar systematic temperature errors were observed in retrieval methods using SIRS-A data [Fritz, Wark, et al., 1972] and VTPR data [Jastrow and Halem, 1973]. In both references, the errors were compensated for by adjusting, or "tuning", the transmittance values channel by channel until the differences between retrieved profiles and the verifying radiosonde profiles at island stations used as check-profiles were brought within acceptable limits. However in this study, without radiosondes to use as check-profiles it was not possible to attempt "tuning" in a sophisticated manner.

Nevertheless, to maintain quality control in the retrieved profiles it was necessary to eliminate the superadiabatic lapse rates between 1000 and 900 mbs. This could have been done more elegantly, however reducing transmittance values for channel five by a factor of .95 and values for channel six by a factor of .90 proved to be

sufficient for the present purpose. Typical examples of retrieved profiles before and after the "tuning" of the transmittances as just described are depicted in Figs. 6 and 7. Note that the super-adiabatic lapse rates between 1000 and 900 mbs have been corrected by the "tuning" process which will be discussed further in Section V.

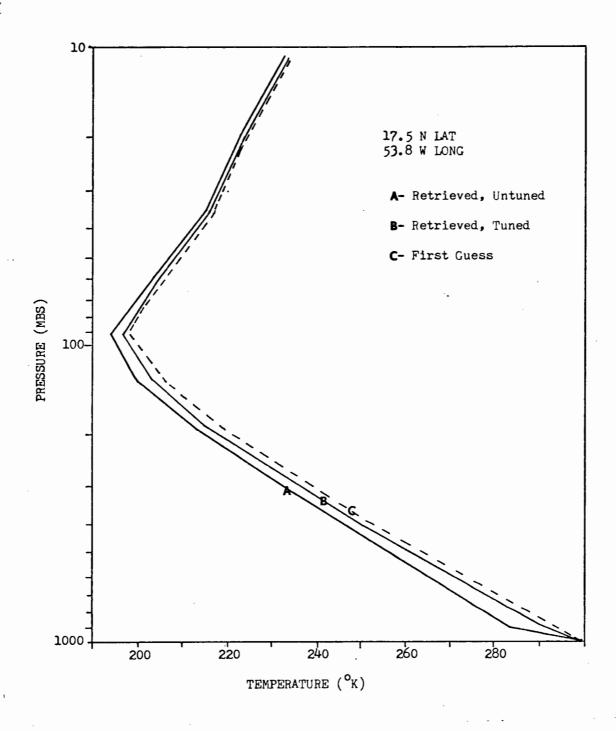


FIG. 6. Low-latitude temperature retrievals using both untuned and tuned transmittances. Note elimination of super-adiabatic lapse rate in lowest 100 mb by the tuned transmittances.

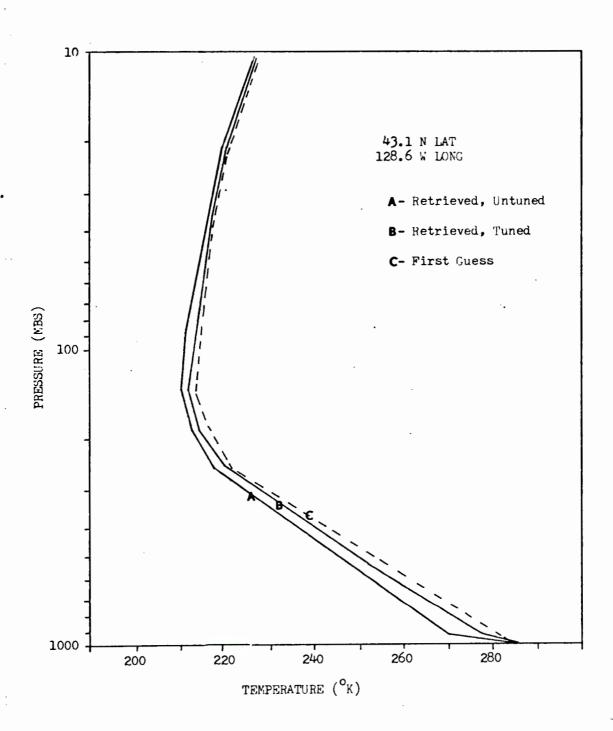


FIG. 7. Mid-latitude temperature retrievals using both untuned and tuned transmittances. Note elimination of superadiabatic lapse rate in lowest 100 mb by the tuned transmittances.

## V. TRANSMITTANCE TUNING

Although there are other possible causes of systematic error in radiance retrieval results, such as improper instrument calibration, the largest single source of such errors is presumably due to uncertainties in the transmittance functions [Drayson, 1971].

Transmittances are computed from theoretical models of absorption band structure, and are uncertain by at least a few percent in each channel [Jastrow and Halem, 1973]. In the case of transmittance functions for the 15  $\mu m$  carbon dioxide channels, much of this uncertainty is due to lack of precise knowledge of the intensities, half-widths, and shapes of the absorption lines of all molecular absorbers contributing to  $\tau_1(p)$ . Such knowledge should also be available for the 14  $\mu m$  band of ozone, and for the 20  $\mu m$  pure rotational band of water vapor [Drayson, 1971]. The ozone band absorbs weakly in the upper stratosphere, while a weakly absorbing edge of the water vapor band is effective in the lower troposphere.

In addition, the transmittance functions are known to be weakly temperature dependent. For a given standard temperature profile the total transmittance of the atmosphere for VTPR carbon dioxide channels can be considered to be the product of the individual transmittance of carbon dioxide, ozone, and water vapor [McMillin, Wark, et al., 1973]:

$$\tau(p) = \tau_{CO_2}(p) \cdot \tau_{O_3}(p) \cdot \tau_{H_2O}(p)$$
 (22)

Since the transmittance values used in this study were carbon dioxide transmittances for a mid-latitude standard atmosphere profile and not subjected to the correction of (22), it was assumed that the standard listings of  $\tau(p)$  were not completely descriptive of the absorption profile for the real atmosphere on the day the VTPR measurements were made. These factors are justification for periodic tuning in general [Jastrow and Halem, 1973], and in particular were considered to be the source of the systematic negative temperature errors in the initial retrieved profiles already discussed. Based on this assumption, the transmittances were "tuned" to correct the super-adiabatic temperature lapse rates between 1000 and 900 mbs.

Since the atmosphere transmits more strongly in certain pressure intervals than others for a given channel, the temperature of a given layer can be adjusted by tuning the transmittances for channels which receive greater radiance contributions from the layer. The selective transmittance of the atmosphere is illustrated in Fig. 8 for the "untuned" transmittances. It is clear from the figure that in the layer 1000 to 900 mb the largest atmospheric transmittance is in channels five and six.

From the quadrature form of the radiative transfer equation (7), for a given channel radiance value,  $I_{\underline{i}}$ , it follows that selectively decreasing the layer weighting values,  $\Delta \tau_{\underline{i}}(K)$  requires that the value of the layer mean Planck values,  $\overline{B}_{\underline{i}}^{N}(K)$ , be increased. An increase in the

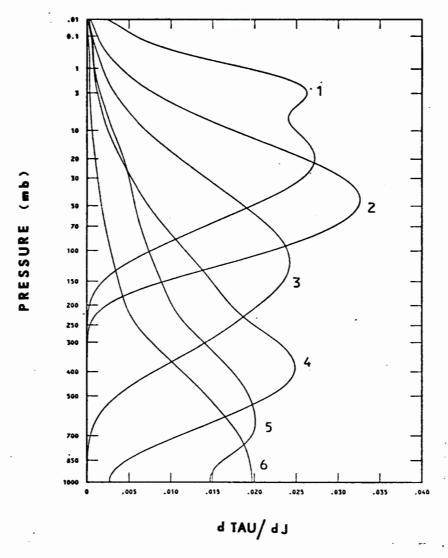


FIG. 8. Weighting functions for VTPR carbon dioxide channels.

layer-mean Planck value will give a higher value for the weighted layer mean Planck for a specific layer evaluated by eq. (15), and that in turn will give a higher layer-mean temperature in accordance with eq. (18).

Increasing the retrieved mean temperature for any layer, K, could therefore be accomplished by decreasing the  $\Delta \tau_{\bf i}(K)$  for channels receiving radiance contributions from the layer.

Since channels five and six most significantly affected the layer 1000 to 900 mbs (as indicated in Fig. 8), increasing the retrieved temperature of this layer by decreasing  $\Delta \tau_5(K)$  and  $\Delta \tau_6(K)$  was feasible.

Rather than decrease transmittance values for channels five and six only in the lowermost layers, a constant fractional decrease of transmittances was applied at all J-levels thereby decreasing  $\Delta \tau_5(K)$  and  $\Delta \tau_6(K)$  for all layers, K. Decreasing the transmittances only in the lowermost layers would have increased the temperature of the layer 1000 to 900 mbs, but it also would have had the effect of changing the shape of the  $\frac{d\tau(J)}{dJ}$  weighting functions for channels five and six. Since a main objective of this thesis was to determine which atmospheric layer thicknesses were best specified by the VTPR "clear column" radiances, a tuning method which did not change the layers in which the  $\frac{d\tau(J)}{dJ}$  curves "peaked" was considered desirable, in the absence of radiosonde check-profiles to justify more complex tuning. The fractional decrease of transmittances at all levels satisfied this condition and had the effect of shifting the  $\frac{d\tau(J)}{dJ}$  curves of Fig. 8 for channels five and six slightly to the left while retaining their original shape.

The selection of transmittance tuning factors, .95 for channel five and .90 for channel six, was strictly empirical. However subsequent to completion of the retrieval computations, it was learned through private communication [Dr. L. M. McMillin, 1973] that NESS had made somewhat similar tuning adjustments for these two channels for the same VTPR data analysis.

## VI. RETRIEVED THICKNESS ANALYSIS

#### A. INDIVIDUAL LAYER SPECIFICATION

Individual layer thicknesses were subjected to stepwise multiple regression analysis to determine which particular atmospheric layer thicknesses were best specified by using the "clear-column" radiance measurements as predictors.

Prior to statistical analysis, retrieved thicknesses were separated into three samples based on the latitudes of their corresponding VTPR scan spots. Latitude bands and number of scan spots for which thicknesses were retrieved are listed in Table 2.

TABLE 2. Scan spot samples for statistical analysis of retrieved thickness values.

Latitude Band	Latitude of Scan Spots	Sample Size
. 1	15 N < Lat < 25 N	82
2	25 N <u>&lt;</u> Lat < 35 N	104
3	35 N <u>&lt;</u> Lat < 45 N	109

## 1. Stepwise Regression Analysis

The Stepwise Regression Analysis Program BIMED 02R [Dixon, 1966] was used to analyze thickness data from the three latitude band samples.

BIMED 02R computes, in a stepwise manner, a sequence of linear regression equations, with one variable added to the regression equation at each step. The variable added is the one that results in the greatest reduction in the previously unexplained sum of squares. This is also the variable which has the highest partial correlation with the dependent variable at the particular step in the analysis of variance. In addition, it is the variable which would have the highest  $F_{\rm L}$ -value when added at step k.

The  $F_k$ -value at each step k is [Crow, et al., 1955]

$$F_k(1,n-k-1) = \frac{\%(C.E.V.k) - \%(C.E.V.k-1)}{\%(U.E.V.k)}$$
 (24)

where

%(C.E.V.k) is the percent cumulative explained variance at step k.

%(U.E.V.k) is the percent unexplained variance at step k.

This study employed a statistical model expressed as

$$\Delta \mathbf{z} = c_0 + c_1 N_1 + c_2 N_2 + c_3 N_3 + c_4 N_4 + c_5 N_5 + c_6 N_6 + c_7 N_7$$
(25)

where  $\Delta z$  is a standard layer thickness, N<sub>1</sub> through N<sub>6</sub> are radiance predictors corresponding to VTPR "clear-column" radiance measurements of channels one through six, and

 $N_7$  is that computed for the window channel (channel eight) from the sea surface temperature by assuming a transmittance of unity [McMillin, Wark, et al., 1973].  $C_0$  through  $C_7$  are regression coefficients computed by BIMED 02R.

variables, or predictands, were all possible layer thicknesses between the pressure levels considered from 1000 to
100 mb. For examination of the lower stratosphere,
predictands were all possible layer thicknesses from 100
to 20 mb. Each dependent variable was subjected to the
BIMED 02R stepwise multiple regression analysis to determine
how well the layer thickness was specified by the independent
variables, the "clear-column" radiances.

Related statistical parameters from the BIMED 02R output included:

- a. multiple R
- b. standard error of estimate, S.E.
- c. mean value,  $\overline{\Delta z}$
- d. standard deviation,  $\sigma$
- e. F<sub>k</sub>-value
- $f. R^2$

The most significant parameters for purposes of this study were  $R^2$  and S.E., which are related as follows:

$$(S.E.)^2 = \sigma^2 [(n-1)/(n-k-1)](1-R^2)$$
 (26)

. where

$$\sigma^2 = \sum_{i=1}^{n} (\Delta z_i - \overline{\Delta z})^2 / (n-1)$$
 (27)

is the variance of a layer thickness for the sample, and

n = sample size

i = sample-element identifier

k = number of predictors selected (k = 7 in this study).

The fractional unexplained variance  $(1-R^2)$ , can be approximated by

$$(1 - R^2) \stackrel{!}{=} \left(\frac{S.E.}{\sigma}\right)^2 \tag{28}$$

since (n-k-1)/(n-1) is close to unity for the sample sizes and number of predictors considered.

The fractional explained variance can then be expressed

$$R^2 = 1 - \left(\frac{S.E.}{\sigma}\right)^2 \tag{29}$$

which becomes the percent explained variance upon multiplying  $R^2$  by 100.

The greater the percent explained variance for a particular layer thickness, the better the thickness is considered to be specified by the radiance predictors.

# 2. Individual Layer Results

Although the specification of individual layers varied somewhat from latitude band to latitude band, in

general, the layer thicknesses best specified by the VTPR
"clear-column" radiances were for relatively large pressure
intervals, where also the lower pressure level was in the
lower troposphere. Layer thicknesses most poorly specified
were for pressure intervals in vicinity of the tropopause.

A total of 99 different layer thicknesses were examined between 1000 and 20 mb for each latitude band. Rather than tabulate the statistical parameters for all 99 layers, a representative group of layer thickness results has been extracted which illustrates the general trends in thickness specifications.

Statistical parameters for a set of sequential layers between 1000 and 20 mb for all three latitude bands are listed in Table 3. Note that layers in vicinity of the tropopause, that is those between 150 and 90 mb, were not as well specified in terms of fractional explained variance as the layers above and below. This is indicative of "noise" in the retrieved mean temperatures near the tropopause associated with the reversal of temperature gradient across the interface. Both Tables 3 and 4 show that the explained variance of the thickness of layers in vicinity of the tropopause (e.g. 150-100, 100-90 mb) increases with increased latitude. Statistically this was a result of the larger standard deviations in the thicknesses of layers 150 to 100 mb and 100 to 90 mb as one progresses into midlatitude bands. Give a larger  $\sigma^2$ , the variable  $\Delta z$  is more predictable in terms of layer-mean temperature and/or radiances.

TABLE 3. Statistical parameters of some representative sequential layers.

	Ban	d 1	Ban	d 2	Ban	<b>d</b> 3
Pressure Layer (mb)	S.E. (gpm)	R <sup>2</sup>	S.E. (gpm)	$R^2$	S.E. (gpm)	R <sup>2</sup>
30–20	1.670	•9779	1.487	.9856	1.661	.9865
50-30	2.123	.9810	2.981	.9693	2.940	.9758
70–50	6.175	.8110	5.731	.8622 ·	5.474	.8982
90-70	7.966	.6627	6.222	.8083	5.650	.8662
100-90	3.792	.6052	2.914	.7847	2.623	.8559
150-100	9.352	.6076	8.907	.7487	8.558	.8622
200–150	1.594	.9317	3.686	.8021	3.752	.9283
250–200	0.489	.9944	0.844	.9876	0.802	.9954
300–250	0.748	.9863	0.514	.9956	0.532	.9978
400-300	1.572	.9803	1.436	.9897	1.504	.9945
500–400	1.255	.9821	1.469	.9853~	1.531	.9921
600–500	0.957	.9861	1.415	.9816	1.487	.9901
700–600	2.532	.9777	1.403	.9769	1.467	<b>.9</b> 877
1000-700	2.532	•9797	2.929	.9813	2.960	.9894

TABLE 4. Statistical parameters of some layers of special interest

_		nd l	Ban	d 2	Band	1 3
Pressure Layer (mb)	S.E. (gpm)	R <sup>2</sup>	S.E. (gpm)	R <sup>2</sup>	S.E. (gpm)	R <sup>2</sup>
300-150	0.384	•9995	4.053	.9638	4.072	.9862
400-150	1.230	.9982	2.826	•9932	2.724	•9974
500-150	2.496	•9957	1.778.	.9985	. 1.594	•9995
600-150	3.452	.9944	1.505	•9993	1.462	.9997
700–150	4.540	•9927	2.338	.9987	2.446	.9994
150-100	9.352	.6076	8.907	.7487	8.558	.8622
200-100	11.006	.6565	12.603	.7383	12.306	.8813
250-100	10.520	.7481	13.319	.7679	13.084	.9105
100-90	3.792	.6052	2.914	.7847	2.623	.8559
100-70	11.704	.6449	9.136	.8008	8.274	.8628
1000-300	7.399	.9808	8.668	.9833	8.939	.9909
1000-150	7.090	.9989	5.011	.9892	5.156	.9982
1000-100	2.390	.9988	4.351	.9972	4.245	.9989
150-20	27.430	.8435	26.164	.8787	24.616	.9144
100-20	17.445	.9024	16.230	.9239	15.0312	.9446

The sequential layers best specified were those between 600 and 200 mb. These layers are below the highly variable tropopause region, and furthermore are in a pressure interval where significant transmittance of radiance in channels three through six occurs (see Fig. 8). Presumably, having significant radiance contributions in a given layer by several VTPR channels improves specification of the layer thickness by the radiance measurements.

Statistical parameters for thicknesses of layers of special interest are listed in Table 4. These include the layers generally best specified, the layers generally most poorly specified, the 1000 to 300 mb layer thickness currently used in the FNWC vertical-analysis scheme [Holl et al., 1964], and the defined "tropospheres" and "stratospheres" to be discussed in Sections VI (B) and (C). As in the case of sequential layers already discussed, the most poorly specified of all pressure intervals in Table 4 were shallow and had boundaries near the tropopause. The best specified layers were those in the mid-troposphere where significant radiance contributions from at least four of the CO<sub>2</sub> channels occur.

#### B. TROPOSPHERIC SPECIFICATION BY SUB-LAYERS

Results of the statistical regression analysis applied to individual layers were next used to examine various combinations of the sequential sub-layers in the troposphere to determine which of the combinations could best specify the full tropospheric thickness.

## 1. Method of Analysis

To compare the results of different tropospheric thickness combinations, a normalized root mean square error index was defined. The error index, E.I., for a two-layer combination, for example, was

$$E.I. = \frac{\sqrt{S.E._1^2 + S.E._2^2}}{\overline{\Delta z}_{trop}}$$
 (30)

where S.E.<sub>1</sub> and S.E.<sub>2</sub> are the standard errors of estimate for the individual layers, and  $\overline{\Delta z}_{trop}$  is the mean thickness of the defined troposphere. Similarly, the E.I. for a three-layer combination was

E.I. = 
$$\frac{\sqrt{S.E._{1}^{2} + S.E._{2}^{2} + S.E._{3}^{2}}}{\overline{\Delta z}_{trop}}$$
 (31)

Weighting of standard errors when computing the E.I. for a particular layer combination was not required since the standard error of thickness reflects variance in the retrieved mean temperatures, and weighting of mean temperatures by  $\ln(p_1/p_2)$  factors was implicit in conversion of mean temperatures to thicknesses by the hypsometric equation.

This error index expresses combined layer errors in a root mean square (RMS) sense, and the smaller the E.I. value the better the combination of layers specifies the total thickness.

Specification of the tropospheric thickness was considered by defining the troposphere in one series of tests to be the layer 1000 to 100 mb, and in a second series of tests to be the layer 1000 to 150 mb. In each case, best specification of the "troposphere" was sought by comparing error-index values for various sub-layer combinations within the troposphere.

To form combinations of layers, the concept of a "sliding layer" was adopted. The "sliding layer" was first tied down to the 1000 mb level and the RMS error, as expressed by the error index, was computed for the resulting two-layer combination. The sliding layer was then stepped to a higher level and the RMS error for the resulting three-layer combination calculated. The procedure of stepping the sliding layer to higher levels to form new combinations was continued until a final two-layer combination was formed with the top of the sliding layer being coincident with the defined tropopause level (150 or 100 mb). The pressure interval of the sliding layer was then increased and the procedure repeated. "Sliding layers" of 300, 400, 500, 600, 700, and 800 mb were employed. Fig. 9 illustrates the combinations of sublayers formed by the 400 mb sliding layer in the 1000 to 100 mb troposphere analysis.

Error index values from all combinations of layers formed by the above procedure were compared to determine which particular combination gave the minimum RMS error for specification of the full tropospheric thickness. In

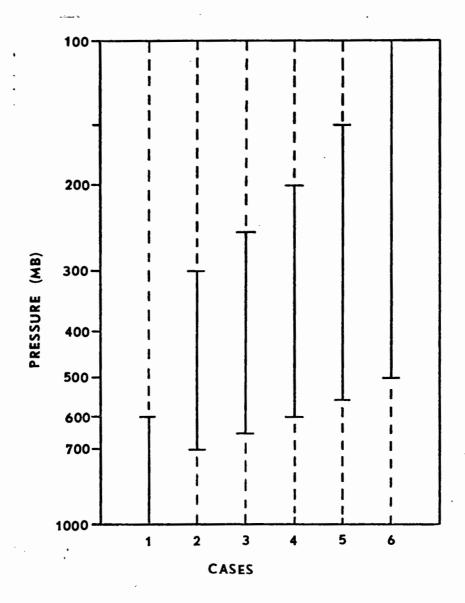


FIG. 9. All combinations, or cases, of tropospheric sub-layers formed by the 400 mb "sliding layer" with the tropopause defined at 100 mb. In each case, the combined sub-layers span the pressure interval 1000-100 mb and the location of the sliding layer is indicated by the solid line.

addition, the combination layer results were compared against the E.I. values for the "tropospheres" considered as single layers, 1000 to 150 mb and 1000 to 100 mb.

## 2. Troposphere Results

In all three latitude bands, defining the troposphere to be the pressure interval 1000 to 150 mb, rather than 1000 to 100 mb, resulted in generally lower E.I. values for the layer combinations considered. This result can be attributed to the fact that the single layer 150 to 100 mb was the most poorly specified of all troposphere layers due to variability induced by the tropopause in this layer. Therefore defining the model tropopause at 150 mb gave smaller combined error index values by eliminating errors of the 150 to 100 mb layer.

In latitude bands two and three, the maximum RMS error for both two and three layer combinations occurred when the lower limit of the top layer of a combination was at 250 mb. In latitude band one this maximum occurred when the lower limit of the top layer was at 200 mb. This result may also be explained in terms of influence of the tropopause, since the tropopause is generally higher in lower latitudes [Haltiner and Martin, 1957], and therefore its detrimental effect on the standard error of estimation is greater for the layer 200 to 100 mb than for the layer 250 to 100 mb as in the higher latitude bands. Standard error values for these layers listed in Table 4 illustrate this relationship.

The general trend in the troposphere was for the combination error index values to be lowest for certain two-layer combinations topped at 150 mb with a dividing level near 650 mb (Table 5). Table 5 lists in sequence the layer-combinations found here to be most effective in giving minimum error index values. Examination of Figs. 10, 11, 12 show that other layer choices give rise to larger RMS errors of specification in the troposphere. No attempt has been made in this study to ascribe statistical significance to any one layer choice because of the limited period of the test. However, the general nature of the "best" results in each band seem to resemble the tropospheric stratification already described.

When the RMS errors for layer combinations were compared against those for the "tropospheres" taken as single layers, it was found that the layer 1000 to 150 mb could be better specified by certain combinations of sublayers than when considered as a single layer. The layer 1000 to 100 mb, however, had the minimum of all RMS errors found when considered as a single layer. This evidently was due to the fact that the relatively poor specification in the 150 to 100 mb layer became insignificant in comparison to the thickness-specification of the layer 1000 to 100 mb. Error index values for the two defined troposphere layers considered as single layers are listed by latitude band in Table 6, and these may be compared with the best combination results in Table 5. Note that the best

TABLE 5. Tropospheric sub-layer combinations having the smallest error index values.

Pressure Layers	(mb)	Band 1	E.I.
1000-700, 700-400,	400-150		.000315
1000-700, 700-100			.000341
1000-550, 550-150			.000370
		.Band 2	
1000-700, 700-150	•		.000275
1000-650, 650-150			.000308
1000-600, 600-150			.000338
•		Band 3	
1000-700, 700-150			.000284
1000-650, 650-150		•	.000316
1000-600, 600-150			.000345

specification of the troposphere in each latitude band was obtained by considering the single alyer 1000 to 100 mb.

TABLE 6. Error index values for the defined "tropospheres" considered as single layers.

"Troposphere" (mb)	Band 1 E.I.	Band 2 E.I.	Band 3 E.I.
1000 to 100	.000148	.000271	.000265
1000 to 150	.000515	.000364	.000382

Error index values for all tropospheric layer combinations considered, as well as values for the defined tropospheres taken as single layers, are graphically depicted in Figs. 10, 11, and 12. For each latitude band, graphs (a), (b), (c) depict E.I. values for layer combinations formed by "sliding layers" of 300, 400, and 500 mb. The (d) graph for each band depicts E.I. values for all other combinations considered, that is combinations resulting from sliding layers of 600, 700, and 800 mb, as well as error index values for the single layer troposphere cases.

Note that except for the two-layer combination 1000 to 700 and 700 to 100 mb in latitude band one, all other combinations formed by sliding layers had smaller error index values when the tropopause was defined at 150 mb. Also note the shift in maximum E.I. value from

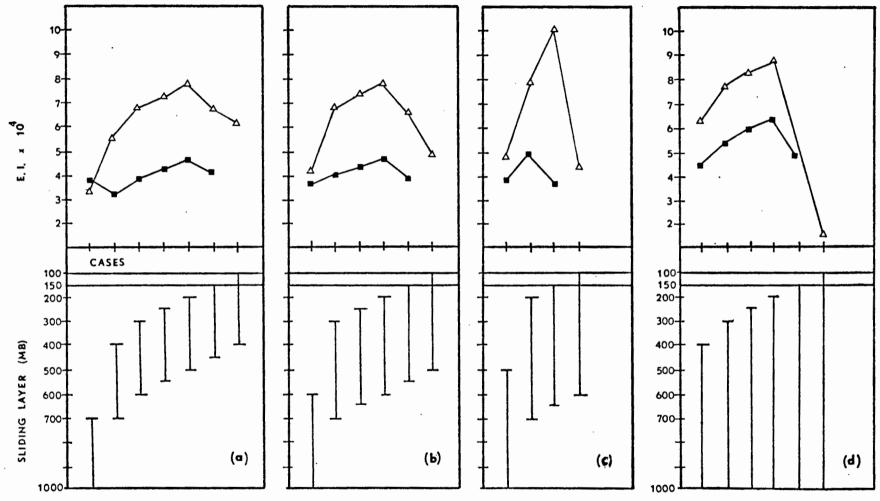


FIG. 10. Cases of tropospheric sub-layer combinations considered and corresponding error index values (15 N to 25 N).

■ E.I. values with tropopause at 150 mb. △ E.I. values with tropopause at 100 mb.

(a) Sub-layer combinations formed by sliding layer of 300 mb, (c) 500 mb, (d) all others.

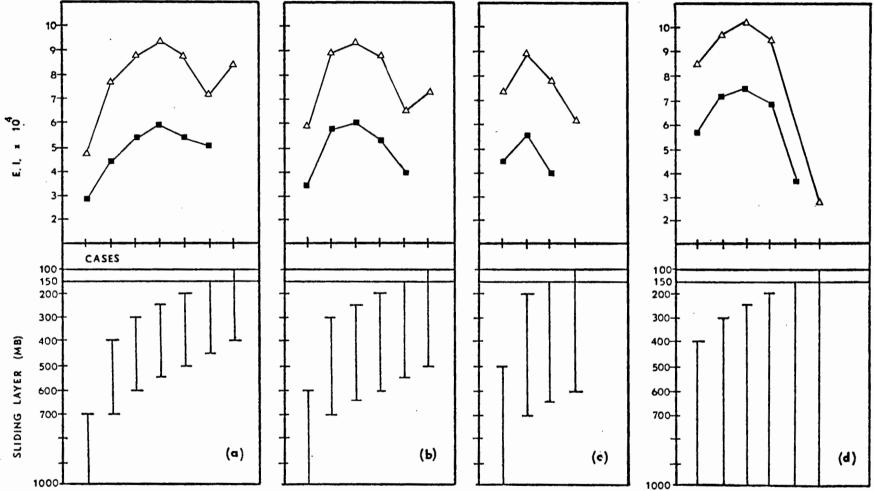
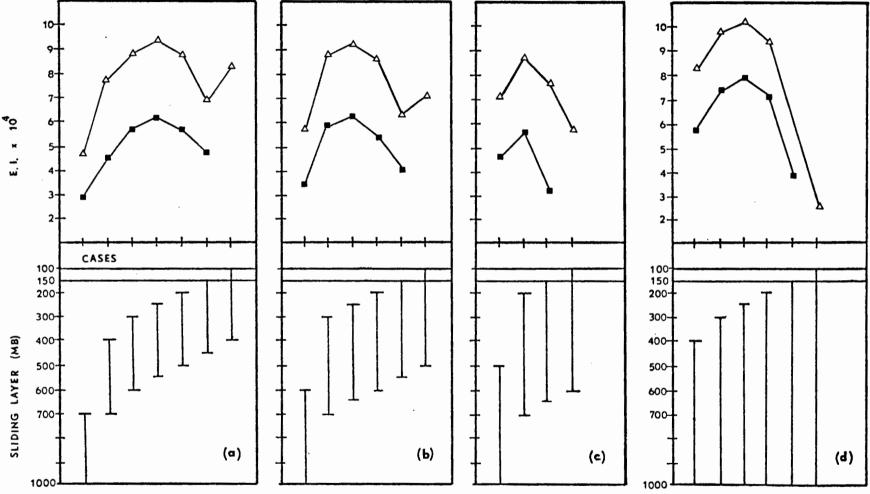


FIG. 11. Cases of tropospheric sub-layer combinations considered and corresponding error index values (25 N to 35 N).

■ E.I. values with tropopause at 150 mb. (a) Sub-layer combinations formed by sliding layer of 300 mb, (b) 400 mb, (c) 500 mb, (d) all others.



Cases of tropospheric sub-layer combinations considered and corresponding error index values (35 N to 45 N).

■ E.I. values with tropopause at 150 mb. (a) Sub-layer combinations formed by sliding layer of 300 mb, (b) 400 mb, (c) 500 mb, (d) all others

combinations with the topmost layer having a lower limit at 200 mb in band one to combinations with the topmost layer having a lower limit at 250 mb in bands two and three.

#### C. STRATOSPHERIC SPECIFICATION BY SUB-LAYERS

Best specification of the stratosphere to 20 mb was sought in a manner similar to that of the troposphere already discussed.

## 1. Method of Analysis

The error index was as defined in specification of the troposphere, except that  $\overline{\Delta z}_{strat}$  was used instead of  $\overline{\Delta z}_{trop}$  in eqs. (30), (31). The stratosphere was considered in one series of tests to be the layer 150 to 20 mb, and in a second series of tests to be the layer 100 to 20 mb. "Sliding layers" of 30, 40, 50, 60, 70, and 80 mb were used to form layer combinations. Figure 13 illustrates the layer combinations formed by the 30 mb sliding layer between 150 and 20 mb.

## 2. Stratosphere Results

Although the results do not afford as sharp a specification as for the troposphere, a general result was that the 100 to 20 mb "stratosphere" had lower RMS errors for layer combinations than the 150 to 20 mb case. This again can be attributed to the influence of the tropopause in the layer 150 to 100 mb. By defining the tropopause to be at 100 mb the poorer specification of the layer 150 to 100 mb was eliminated.

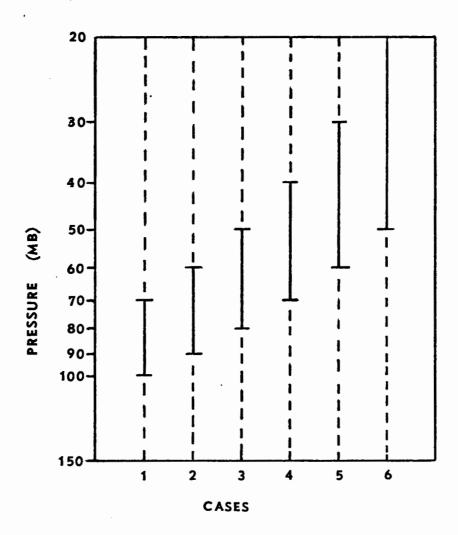


FIG. 13. All combinations, or cases, of stratospheric sub-layers formed by the 30 mb "sliding layer" with the tropopause defined at 150 mb. In each case, the combined sub-layers span the pressure interval 150-20 mb and the location of the sliding layer is indicated by the solid line.

The three best layer combinations in each latitude band are listed sequentially in Table 7 with their corresponding error index values. There was not a unique set of combinations that proved to be most effective for all latitudes, but a general trend for the error index of best combinations to decrease with increasing latitude is evident. This apparently was a result of the tropopause being at lower levels for higher latitudes, which tended to reduce the influence of the tropopause on specification of the defined stratosphere above 100 mb at higher latitudes. As in the tropospheric analysis, no attempt has been made to attach statistical significance to the results due to the limited period of the test.

considered as single layers were not as good as for combinations of layers. A decrease in E.I. values when the 150 to 100 mb layer was incorporated into the total thickness was not observed as in the troposphere analysis. This difference was evidently due to the fact that the thickness-error induced by the 150 to 100 mb layer was relatively smaller in comparison to the specification of the full tropospheric thickness, but had greater influence on thickness-error of the stratosphere. Error index values for the two defined stratosphere layers are listed by latitude band in Table 8.

Error index values for all stratospheric layer combinations considered, as well as values for the defined

TABLE 7. Stratospheric sub-layer combinations having the smallest error index values.

Pressure Layers (mb)	Band 1	E.I.
100-90, 90-60, 60-20		.00123
100-80, 80-20		.00130
150-100, 100-70, 70-20		.00131
	Band 2	
100-90, 90-60, 60-20		.00105
100-80, 80-50, 50-20		.00112
100-70, 70-20	•	.00116
	Band 3	
100-70, 70-30, 30-20		.00094
100-80, 80-40, 40-20		.00095
100-90, 90-60, 60-20		.00096

TABLE 8. Error index values for the defined "stratospheres" considered as single layers.

"Stratosphere" (mb)	Band 1 E.I.	Band 2 E.I.	Band 3 E.I.
100 to 20	.00172	.00162	.00148
150 to 20 .	.00220	.00214	.00198

stratospheres taken as single layers, are graphically depicted in Figs. 14, 15, and 16. For each latitude band, graphs (a), (b), (c), (d) depict E.I. values for layer combinations formed by 30, 40, 50, and 60 mb "sliding layers"; (e) graph depicts error index values for the single-layer stratosphere cases in addition to those for the 70 and 80 mb sliding layer combinations.

The graphical display of stratospheric error index values illustrates the fact that defining the tropopause to be at 100 mb gave better E.I. values than when the tropopause was defined at 150 mb and the layer 150 to 100 mb was included in the stratospheric specification. Also note that the E.I. values are an order of magnitude greater than the tropospheric error index values previously examined.

Apparently, the significantly greater E.I. values in the stratosphere are due to standard errors of estimate, S.E., of layers in the stratosphere reflecting uncertainties of explained variance for the tropospheric column below, as well as for the layers of the stratosphere. In the

troposphere, on the other hand, the layer combinations are "tied down" to the surface and tropospheric standard errors reflect only unexplained variance in the troposphere. Essentially, the relatively greater S.E. values of the stratosphere result in correspondingly greater E.I. values, although  $\overline{\Delta z}$  of the denominator is somewhat smaller than for the tropospheric case.

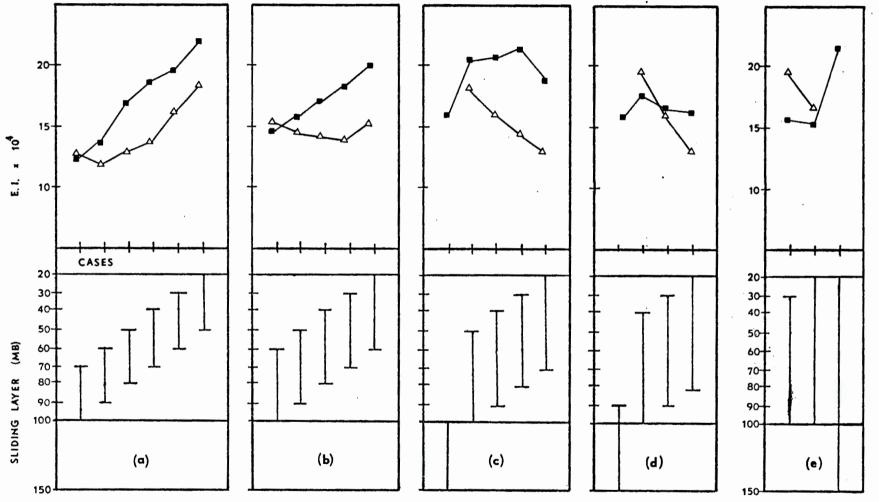


FIG. 14. Cases of stratospheric sub-layer combinations considered and corresponding error index values (15 N to 25 N).

■ E.I. values with tropopause at 150 mb. Δ E.I. values with tropopause at 100 mb.

(a) Sub-layer combinations formed by sliding layer of 30 mb,

(b) 40 mb, (c) 50 mb, (d) 60 mb, (e) all others.

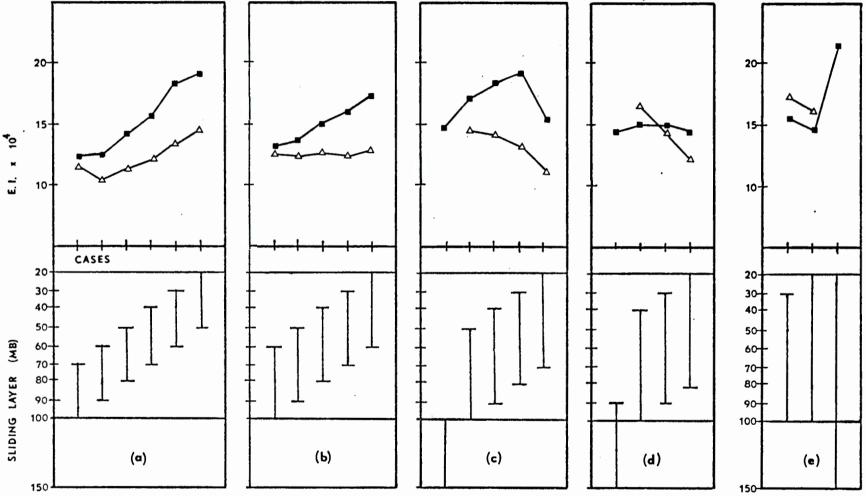


FIG. 15. Cases of stratospheric sub-layer combinations considered and corresponding error index values (25 N to 35 N).

■ E.I. values with tropopause at 150 mb. Δ E.I. values with tropopause at 100 mb.

(a) Sub-layer combinations formed by sliding layer of 30 mb,

(b) 40 mb, (c) 50 mb, (d) 60 mb, (e) all others.

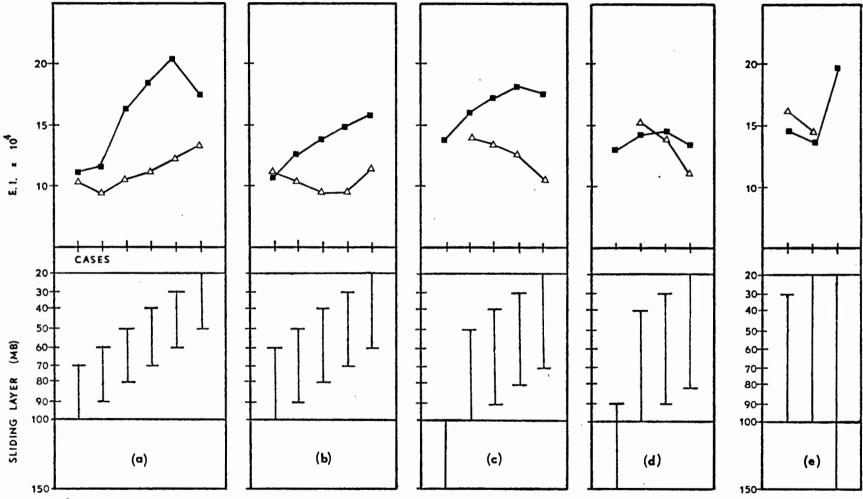


FIG. 16. Cases of stratospheric sub-layer combinations considered and corresponding error index values (35 N to 45 N).

■ E.I. values with tropopause at 150 mb. △ E.I. values with tropopause at 100 mb.

(a) Sub-layer combinations formed by sliding layer of 30 mb, (b) 40 mb, (c) 50 mb, (d) 60 mb, (e) all others.

### VII. CONCLUSIONS

This thesis has demonstrated that atmospheric thicknesses can be retrieved from VTPR "clear-column" carbon dioxide radiances by an improved iterative technique for direct solution of the radiative transfer equation, and that systematic errors in the retrieval results can be corrected by selective adjustment of the transmittances. In addition, the retrieved thicknesses of certain standard layers appear to be better specified than others when the "clear column" radiances are used as regression predictors. Furthermore, it has been found in this case that specification of the full troposphere and stratosphere layers can in some instances be improved by decomposition of the layer into sequential pressure intervals.

It should be noted that the study conducted here bears more heavily on the retrieval technique, and utilizes simulated atmospheric-layer thicknesses which have had "clear-column" radiances built into them. In order to make any final inferences as to choice of an optimum layer thickness, the procedure suggested here would have to be conducted over a period of time and retrieved thicknesses verified against thicknesses computed from corresponding radiosonde soundings.

CLIMATOLOGICAL TEMPERATURE PROFILES USED TO DERIVE FIRST GUESS PROFILES FOR VTPR "CLEAR COLUMN" SCAN SPOTS 15N TO 45N

APPENDIX A

LEVEL	PRESSURE	15N ANN	30N JAN	30N JUL	45N JAN	45N JUL
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The following procedure was used to expand standard 56 level climatological temperature profiles to 100 level profiles on the J scale:

- 1. Where pressures of J-levels were equal to pressures of standard levels, the temperature at the J-level was set equal to the temperature at the standard level.
- 2. Where pressures of J-levels were not equal to pressures of standard levels, temperatures at the J-levels were computed by interpolation as follows:

$$T(J) = T_{STD}(M-1) + \left[\frac{T_{STD}(M) - T_{STD}(M-1)}{P_{STD}(M) - P_{STD}(M-1)}\right][P(J) - P_{STD}(M-1)]$$

where

T(J) = desired temperature at the J-level

P(J) = pressure at the J-level

T<sub>STD</sub>(M-1)= known temperature at nearest standard pressure level above the J-level

T<sub>STD</sub>(M)= known temperature at nearest standard pressure level below the J-level

 $P_{STD}^{(M-1)}$  = pressures at nearest standard levels above and  $P_{STD}^{(M)}$  below the J-level.

## APPENDIX B

### UNTUNED TRANSMITTANCES AND CORRESPONDING STANDARD TEMPERATURES AND J-SCALE PRESSURES FOR THE SIX VTPR "CLEAR COLUMN" CARBON DIOXIDE CHANNELS

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ပ ပ		INTI N L	ER F	PO	LA RC	T	E J	F	OR N	Α	n E	) V P	R I J l	J L	**	S S T	T	D I	T PR	E:	M F	ı L	P	RC S	F	ΙL	. E	S	Δ	T	3	0	A	ND	4	5
	3	DO : TREI TREI CON	F(3	3 ,	J)	5 = =	6 (T	R	E F	(	2, 3,	J	) <del>1</del>	- T	R I	EF EF	(	4 5	, J	)	) = ) =	* O		5												
CCC		CCM	PU1	ΓΕ	Р	R	ES	S	F	0	R	1	00	)	L	ΕV	Ε	L:	S	Αl	N	)	T	ΛU	Ε	1	R	14	v S	Μ	ΙΤ	T#	N	CE	S	
	4	P(I TAU TAU CON	) = ( (5) (6)	, İ	01 ) = )=	*	(1	5	华丁	Δ	116	5	. 1	1	•	26	0	8	7 ช	3	6	* (	*	3.	5											
CCC		COM	PUT	ГΕ	L	A	ΥE	R	T	R	AΝ	I S	M ]	T	T	ΔN	С	E:	S																	
-	56	DO CONCORN	5 '	T =	7.	1	03 U ( 03	, ,	6 , I	- T	3 ) (N	· -	T /	U =	( T	N, AU	I	+: N	3)	. } -		ΓΔ	ιU	(!)	,	4)										

```
C
        WRITE (6,56) WRITE (6,55) (DT(N,103),N=1,6),((DT(N,1),N=1,6),I=7,97
       1,6)
CCC
         READ LAT, LONG, RADIANCES, SEA SFC TEMP FOR SCAN SPCT
      7 READ (5,40,END=35) ALAT,ALONG,(RADCBS(N),N=1,6),TSFC
CCC
        CCNVERT LAT, LONG, RADIANCES TO CORRECT DIMENSIONS
         LATREF= 30
        ALAT=ALAT + 0.1-90.0
         ALONG=ALONG*0.1
        DO 8 N=1,6
RADOBS(N)=RADCBS(N)*0.05
      8 CONTINUE
        INTERPOLATE FOR 56 LEVEL PROFILE AT LAT OF SCAN SPOT FROM "APRIL" STD PROFILES N AND S OF SCAN SPOT
        IF (ALAT.LT.LATREF) LATREF=15
FACLAT=(ALAT-LATREF)/15.0
LAT=LATREF/15
        LATP1=LAT+1
        DO 9 J=1,56
TSTD(J)=FACLAT*(TREF(LATP1,J)-TREF(LAT,J))+TREF(LAT,J)
IF (ALAT.EQ.LATREF) TSTD(J)=TREF(LAT,J)
      9 CONTINUE
C
C
C
         SET 1000MB TEMP EQ TO OBSERVED SEA SFC TEMP
        TSTD(56)=TSFC*0.2+269.9
CCC
        EXPAND 56 LEVEL PROFILE FOR SCAN SPOT TO 100 LEVELS
        DO 12 I=1,100
STEP DOWN THRU 56 LEVEL PROFILE UNTIL PRESS EQ OR GT:
CURRENT 100 LEVEL PRESS
C
        DO 10 J=1,56
IF PRESS EQ SET TEMPS EQ, IF PRESS GT INTERPOLATE FOR
C
         TEMP
        iF (PSTD(J).EQ.P(I)) T(I)=TSTD(J)
IF (PSTD(J).GT.P(I)) GC TO 11
    10 CONTINUE
        BEGIN INTERPOLATION ROUTINE
    11
        K=J-1
        RATIO=(TSTD(J<u>)</u>-TSTD(K))/(PSTD(J)-PSTD(K))
        PDIFF=P(I)-PSTD(K)
T(I)=TSTD(K)+(RATIC*PDIFF)
C
        END INTERPOLATION ROUTINE
    12 CONTINUE
C
        WRITE (6,50) ALAT, ALONG
        WRITE (6,51)
0000
        CCMPUTE PRESS AND TEMP OF J=2.5 LEVEL TO COMPLETE FIRST GUESS PROFILE FOR SCAN SPOT
        P(103) = 0.01 \( (1.0 + (1.5) \times 0.26087836 ) \times 3.5 \)
T(103) = T(2) + ((T(3) - T(2)) / (P(3) - P(2)) ) \( (P(103) - P(2)) )
C
        DO 13 J=1,25
J25=J+25
J50=J+50
       J75=J+75
WRITE (6,49) (J,P(J),T(J),J25,P(J25),T(J25),J50,P(J50)
1,T(J50),J75,P(J75),T(J75))
CONTINUE
    13
C
```

```
COMPUTE PLANCK VALUES FOR ALL CHANNELS AND REQUIRED PRESS LEVELS USING FIRST GUESS PROFILE
          DO 15 N=1,6
     DO 14 I=1,103,3

B(N,I)=(C1*CHAN(N)**3)/(EXP(C2*CHAN(N)/T(I))-1)

TGUESS(I)=T(I)

14 CONTINUE
     15 CCNTINUE
          COMPUTE LAYER MEAN PLANCK VALUES
          DO 17 I=7,103,6

DO 16 N=1,6

BMEAN(N,I)=F1*B(N,I-3)+F2*B(N,I)+F1*B(N,I+3)

IF (I.EQ.103) BMEAN(N,I)=F1*B(N,1)+F2*B(N,103)+F1*B(N,
     16 CCNTINUE
     17 CONTINUE
          CALCULATE RADIANCES FOR FIRST GUESS PROFILE
         DC 19 N=1,6

RADSFC(N)=8(N,100)*TAU(N,100)

RADATM(N)=0.0

DC 18 I=7,103,6

RAD=BMEAN(:N,I)*DT(N,I)

RADATM(N)=RADATM(N)+RAD
     18 CONTINUE
          RADCOM(N) = RADATM(N) + RADSFC(N)
          RCIFF(N)=RADOBS(N)-RADCCM(N)
RDIFFA(N)=ABS(RADOBS(N)-RADCOM(N))/RADOBS(N)
     19 CONTINUE
          COMPUTE REF WAVE NO FOR EACH LAYER USING WEIGHTED LAYER MEAN PLANCK VALUES AND BAILEY ITERATION METHOD
          DC 22 I=7,103,6
KCUNT=0
          SUMBDT = 0.0
          SUMDT = 0.0
          DC 20 N=1,6
BDT=BMEAN(N,I)*DT(N,I)
          SUMBOT = SUMBOT+BOT
     SUMDT=SUMDT+DT(N, I).
20 CCNTINUE
          BREF=SUMBDT/SUMDT
         BEGIN BAILEY ITERATIVE TECHNIQUE

X1=665.0+I

IF (I.EQ.103) X1=670.0

TEMP=F1*T(I-3)+F2*T(I)+F1*T(I+3)

IF (I.EQ.103) TEMP=F1*T(1)+F2*T(103)+F1*T(4)

X2=X1-(F(X1)/(DF(X1)-(F(X1)*DDF(X1)/(2.*DF(X1)))))

DIFF=ABS(X2-X1)

V1-Y2
          X1 = X2
          ŔĈUÑŤ=KEUNT+1
          IF (DIFF.GT.EPSREF) GO TO 21
AVCHAN(I)=X2
END BAILEY ITERATIVE TECHNIQUE
     22 CCNTINUE
C
                    (6,53)
(6,54)
(6,47)
          WRITE
                                TWO5, (I, I=7, 97, 6)
                                 AVCHAN(103), (AVCHAN(I), I=7, 97, 6)
          WRITE
          WRITE
WRITE
WRITE
                                 ALAT, ALONG
(RADOES(N), N=1,6)
                     (6,46)
                     (6,43)
                    (6,46)
                                 (RADCOM(N), N=1,6)
          WRITE
          COMPARE OBSERVED AND CALCULATED RADIANCES AND ADJUST
          LAYER MEAN PLANCK VALUES
```

```
C
         KCUNT=0
        KCUNT=KCUNT+1
DO 25 N=1,6
RDIFF(N)=RADOBS(N)-RADCOM(N)
    23
         RDIFFA(N)=ABS(RDIFF(N)/RADOBS(N))
    DG 24 I=7,103,6

BMEAN(N,I)=BMEAN(N,I)+RDIFF(N)

24 CONTINUE

25 CONTINUE
         CALCULATE RADIANCES USING ADJUSTED LAYER MEAN PLANCKS
        DC 27 N=1,6

RADATM(N)=0.0

DG 26 I=7,103,6

RAD=BMEAN(N,I)*DT(N,I)
         RADATM(N)=RADATM(N)+RAD
        CONTINUE
         RADCOM(N)=RADATM(N)+RADSEC(N)
    27 CONTINUE
C
                 (6,44)
         WRITE
         WRITE
                            (RDIFF(N), N=1,6)
(RDIFFA(N), N=1,6)
                  (6,46)
                  (6,46)
                  (6,45)
                            KOUNT
         WRITE
         WRITE
                           (RADCOM(N), N=1,6)
                  (6,46)
         CCMPARE DIFFERENCE BETWEEN CBSERVED AND CALCULATED RADIANCES AGAINST CONVERGENCE CRITERION
         DC 28 N=1,6
RDIFF(N)=RADOBS(N)-RADCOM(N)
    RDIFFA(N)=ABS(RADOBS(N)-RADCOM(N))/RADOBS(N)
IF (KOUNT.EQ.NTEN) GO TO 29
IF (RDIFFA(N).GT.EPSRAD) GO TO 23
28 CONTINUE
C
    GO TO 30
29 WRITE (6,52)
C
        WRITE
WRITE
    30
                 (6,44)
                            (RDIFF(N), N=1, 6)
                  (6,46)
                            (RDIFFA(N), N=1,6)
         WRITE (6,46)
         COMPUTE LAYER MEAN TEMPS FROM FINAL LAYER MEAN PLANCKS
         DO 32 I=7,103,6
SUMDT=0.0
         SUMBDT=0.0
DC 31 N=1,6
BDT=BMEAN(N,I)*DT(N,I)
SUMBDT=SUMBDT+BDT
         SUMDT = SUMDT+DT (N, I)
         CCNTINUE
         BREF=SUMBDT/SUMDT
T(I)=(C2*AVCHAN(I))/ALOG((C1*AVCHAN(I)**3+BREF)/BREF)
    32 CONTINUE
CCC
         INTERPOLATE FOR TEMPS AT INTERMEDIATE LEVELS
        DC 33 I=10,94,6
T(I)=(T(I-3)+T(I+3))*0.5
CONTINUE
         T(4)=(T(103)+T(7))*0.5
         T(1) = T(103) - (T(4) - T(103))
         COMPUTE DIFFERENCE BETWEEN FIRST GUESS AND RETRIEVED TEMP PROFILES
         DO 34 I=1,100,3
TDIFF(I)=T(I)-TGUESS(I)
```

```
34 CONTINUE
C
                              WRITE (6,48) ALAT, ALONG, KOUNT
                              WRITE (6,42)
                              WRITE (6,41) (I,P(I),T(I),TDIFF(I),I=1,100,3)
CCC
                              COMPUTE RADIANCE FOR WINDOW CHANNEL FROM SEA SFC TEMP
                              RADOBS(7) = (C1*835.**3)/(EXP(C2*835./TSTD(56))-1)
C
                              IF ((ALAT.GE.LAT25).AND.(ALAT.LT.LAT35)) LBAND=2
IF ((ALAT.GE.LAT35).AND.(ALAT.LT.LAT45)) LBAND=3
                              PUNCH CARDS FOR THICKNESS PROGRAMS
                        WRITE (7,57) LBAND, ALAT, ALONG, T(1), T(103), (T(I), I=7,971,6), T(100), (RADOBS(N), N=1,7)
C
                              GC TO 7
C
               35 STOP
 C
                                                               (6F6.1)
(8F10.8)
(8F10.4)
                36 FORMAT
                          FORMAT
FORMAT
                37
                                                               (10F8.3)
(9F7.1)
(10F8.3)
(9F7.1)
(10F8.3)
(10F8
                           FORMAT
                39
               40 FORMAT
41 FORMAT
                          FORMAT
                                                              ('O',/,TIO,'CCMPUTED RADIANCES FROM FIRST GUESS
('',TIO,'RADIANCE RESIDUALS')
('O',TIO,'COMPUTED RADIANCES AT ITERATION NO.',
                1NITIAL'
43 FORMAT
                           FORMAT
FORMAT
                          FORMAT
FERMAT
                                                                (1)
                                                                ('',T10,6F12.6)
('0',T10,'OBSERVED RADIANCES',5X,'N. LAT',F7.1,
                46
                1,F7.1)
48 FORMAT
                48 FORMAT ('0',///,T10,'RETRIEVED TEMP PROFILE',5X,'N. LA
1W. LONG', F7.1, T80,'ND OF ITERATIONS NEEDED', I7,/)
49 FORMAT ('',4(8X,I3,4X,F7.3,4X,F7.3))
50 FORMAT ('0',/,T10,'STD TEMP PROFILE',5X,'N. LAT',F7.1,
               1,F7.1)
51 FCRMAT
52 FORMAT
               1,F7.1)
51 FCRMAT ('O',4(8X,'LEVEL',4X,'PRESS',4X,'TEMP',3X))
52 FORMAT ('O',/,T10,'ITERATIONS TERMINATED AT 10, FRACTI
1E, CBSERVED - COMPUTED RADIANCES, FCLLOWS')
53 FORMAT ('O',/,T10,'LEVEL NO',T25,F6.1,1616)
54 FORMAT ('',T10,'REF WAVE NO',T25,17F6.1)
55 FORMAT ('',T10,6F12.5)
56 FORMAT ('O',//,TI0,'LAYER TRANSMITTANCES',/)
57 FORMAT (II,9X,10F7.2/(3X,11F7.2)/(7F9.2))
C
                               END
```

STD TEMP	PROFIL	E N. LA	T 19.2	W. LONG	63.0						
LEVEL 123 45 67 8 9 0 1 1 1 2 3 4 5 6 7 8 9 1 1 1 2 3 4 5 6 7 8 9 1 1 2 3 4 5 6 7 8 9 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2	PR 600333626000000000000000000000000000000	TEMP 8 78 184 8 78 184 8 78 184 8 78 184 8 78 182 8 78 18	F67850173456789C173456789OF6785017333333333444444444456	PR E 5722 113.17565 114.7565 122.146.53 120.6276	TEMP 232.3 73 230.680 228.84C 225.49C 225.49C 225.49C 221.3154 218.9166 217.6877 211.43.278 211.43.278 211.43.278 211.43.278 201.376 2	L 1955555555566666666667777777777777777777	\$\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\	TEMP 4400 . 4980 . 200 . 4980 . 202 . 202 . 204 . 2040 . 2040 . 205 . 8794 . 206 . 8794 . 2011 . 615 . 4777 . 2011 . 615 . 4777 . 2011 . 816 . 4900 . 2011	L 777785012345678901234567890 12345678901234567890 1	PR E \$ \$ 392.47	TEMP 8615 252-86-694 2538-442 2558-442 2562-4375 258-442 2664-3389 2772-6-087 2876-887 2878-887 2885-749 2885-749 2885-749 2885-749 2885-749 2885-749 2885-749 2885-749 2885-749 2885-749
LEVEL NO	NC	2.5	7 13 1 2 6 <b>75.</b> 1 676.	9 25 8 679.4 68	31 1.0 682.7 685	43 43 690.1 69	55 61 6.4 704.0	709.9 714.5 7	79 85 19.3 723.8	728.9 743	97
CBSERVED 54.44				W. LONG 59.399994		98.059991				12017 133	• /
55-40	)A 097	43.703049	RST GUESS PR 43.967941	OFILE 65.803513	82.943344	99.306931					
FADÍANCE -0.95 0.01	7560	0.646942	-2.017944 0.048104	-6.403519 0.107803	-2.793350 0.034852	-1.206940 0.012303					
CCMPUTED 54.45 RADIANCE	RADIAN 7306 RESIGU	CES AT ITER 44.345045 ALS	ATION NO. 41.951385	59.442535	80.642242	98.669693					
54.45 RADIANCE -0.00 0.00	7324 00135	0.000946 0.000021	-0.001389 0.000033	-0.042542 0.000716	-0.492249 0.006142	-0.569702 0.005807					
CCMPUTED 54.45 RADIANCE	0073	CES AT ITER 44.349991 ALS	ATION NG. 41.949997	59 <b>.</b> 40026 <b>9</b>		98.368896					
-0.00	00092	0.0	0.0	-0.000275 0.000005	-0.086761 0.001CE2	-0.268906 0.0C2741		•			
CCMPUTED 54.44 . RADIANCE	RADIAN 19966 RESIDU	CES AT ITER 44.349991 ALS	ATION NO. 41.949997	59.399994	· 80.1652 <b>6</b> 8	98.226898					
0.00	00000	0.0	0.0	0.0	-0.015274 0.000191	-0.126907 0.001294					
CCMPUTED 54.44 RACIANCE	RADIAN 19982 RESIDU	CES AT ITER 44.349991 ALS	ATION NO. 41.949997		80.152695	98.159912		,			
0.0		0.0 0.C	0.0	0.0	-0.0027C1 0.000034	-0.059921 0.000611					

	CCMFUTED RAD 54.449982 RADIANCE RES 0.0	IANCES AT ITE 1DUALS 0.0 0.0 IANCES AT ITE 44.349991 1DUALS 0.0	0:0 ERATION NO. 41.949997 0.0 0.0	59.39\$9 0.0	-0.000488	98.128250 -0.028259 0.000288 98.113342 -0.013351 0.000136		
	CCMPUTEC RAD 54.449982 RADIANCE RES 0.0 0.0	IANCES AT ITE 44.349991 IDUALS 0.0 0.0	ERATION NO. 41.949997 0.0 0.0	59.39 <sup>7</sup> 9	0.000015 0.000000	98.106293 -0.006302 0.006064		
	RETRIEVED TE	MP PROFILE	N. LAT	19.2 k	N. LONG 63.0 PRESS	. NO (	OF ITERATIONS NEEDED	7 RETRIEVED - INITIAL
. :		1470369258147036925814703692581470			0.0100 0.0756 0.2704 0.68363 2.6289 4.4206 1.431504 10.31504 10.31504 120.4203 20.4		179 - 144 2271 - 612 2847 2968 2968 2968 2968 2968 2908 2908 2908 2914 2959 2008 2914 2959 2008 2914 2959 2008 2914 2914 2914 2914 2914 2914 2914 2914	-5.7878620 -6.66116-6.6623 -11.766623 -0.07139211-6.6623 -0.07739711-6.0713911-6.0713-

CCCC	CCNVERSION OF RETRIEVED LAYER MEAN TEMPERATURES TO ATMOSPHERIC THICKNESSES FOR "STANDARD" TROPOSPHERIC LAYERS.
იეიიიიიიიიიიიიი	LT DOUGLAS R. MORAN NAVAL POSTGRADUATE SCHOOL DECEMBER 1973
č	DIMENSION ARRAYS
Č	DIMENSION PSTD(60), TSTD(60), P(19), T(19), DZ(59), Y1
CCC	1(8), Y2(8), RADOBS(7) SET DATA
č	CATA R/28700./,G/980./,LBAND1/0/,LBAND2/0/,LBAND3/0/,N
	10NE/1/,NTWC/2/,NTHREE/3/ RG=R/G
C C C C C C C	READ IN MEAN PRESS LEVELS FROM 100 LEVEL PROFILE AND STD LEVELS FOR 56 LEVEL PROFILE AND ADDITIONAL LEVELS 80, 450, 550, 650, TO COMPLETE "STANDARD" LEVELS
L	READ (5,11) (PSTD(J), J=1,60) READ (5,12) (P(I), I=1,19)
CCCC	READ IN LAT BAND, LAT, LONG, AND TEMPS FOR TOP LAYER MEANS, SFC, FROM TEMP RETRIEVAL PROGRAM
C	1 READ (5,13, END=10) LBAND, ALAT, ALONG, (T(I), I=1,19), (RAD 10BS(N), N=1,7)
C C C	EXPAND 19 LEVEL RETRIEVED PROFILE TO 60 LEVELS
	DO 4 J=1,60 DO 2 I=1,19 IF (P(I).EQ.PSTD(J)) TSTD(J)=T(I) IF (P(I).GT.PSTD(J)) GO TO 3 2 CGNTINUE
C	BEGIN INTERPOLATION ROUTINE 3 K=I-1
	RATID=(T(I)-T(K))/(P(I)-P(K)) PCIFF=PSTD(J)-P(K) TSTD(J)=T(K)+(RATIG*PDIFF)
C	ÉND INTERPOLATION ROUTINE  4 CONTINUE
С	IF (LBAND.EQ.NONE) WRITE (6.14) ALAT.ALONG
С	IF (LBAND.EQ.NONE) WRITE (6,15)
	DC 5 J=1,15 J15=J+15 J30=J+30 J45=J+45 IF (LBAND.EQ.NONE) WRITE (6,16) (J.PSTD(J),TSTD(J),J15 1.PSTD(J15),TSTD(J15),J30,PSTD(J30),TSTD(J30),J45,PSTD 2(J45),TSTD(J45)) 5 CONTINUE
CCC	COMPUTE LAYER THICKNESSES
c	DO 6 I=1,59 DZ(I)=RG*((TSTD(I+1)+TSTD(I))*0.5)*(ALOG(PSTD(I+1)/PST 1C(I))) 6 CONTINUE

```
COMPUTE THICKNESSES ABOVE AND BELOW KEY LEVELS
           DC 9 M=46,56
           K=M-45
Y2(K)=0.0
C
           DO 7 N=46, M
Y2(K)=Y2(K)+DZ(N)
       7 CONTINUE
C
           M1 = M + 1
           Y1(K)=0.0
C
       CO 8 NN=M1,59
Y1(K)=Y1(K)+DZ(NN)
8 CONTINUE
C
           PRESS=PSTD(M1)
C
           IF (LBAND.EQ.NONE) WRITE (6,17) PRESS, Y2(K) IF (LBAND.EQ.NONE) WRITE (6,18) PRESS, Y1(K)
C
       9 CONTINUE
C
           IF (LBAND.EQ.NONE) LBAND1=LBAND1+1
IF (LBAND.EQ.NTWO) LBAND2=LBAND2+1
IF (LBAND.EQ.NTHREE) LBAND3=LBAND3+1
           PUNCH THICKNESS DATA CARDS FOR LATITUDE BANG ONE FOR ANALYSIS BY BIMED 02\text{R}
           IF (LBAND.EQ.NONE) WRITE (7,20) ALAT, ALONG, (RADOES (K),
          1K=1,7)
IF (LBAND.EQ.NONE) WRITE (7,21) ALAT, ALONG, (Y2(K), K=1,
          18)
IF (LBAND.EQ.NONE) WRITE (7,21) ALAT, ALONG, (Y1(K), K=1,
         18)
C
           GO TO 1
C
      10 WRITE (6,19) LBAND1, LBAND2, LBAND3
C
            STCP
                        (8F10.3)
(10F8.3)
(11,9X,10F7.2/(3X,11F7.2)/(7F9.2))
('0',/,T10,'60 LEVEL PROFILE',5X,'N. LAT',F7.1,
      11 FORMAT
12 FORMAT
13 FORMAT
      14 FORMAT
1, F7.1)
15 FORMAT
                         ('0',4(8X,'LEVEL',4X,'PRESS',4X,'TEMP',3X))
('',4(8X,I3,3X,F8.3,3X,F8.3))
('0',T10,'THICKNESS',F9.2,'MB TO 100.00 MB =
          FCRMAT
      16
     16 FCRMAT (' ',4(8X,13,3X,F8.3,3X,F8.3))
17 FORMAT ('0',T10,'THICKNESS',F9.2,' MB TO 100.00 MB =
1TERS')
18 FCRMAT (' ',T10,'THICKNESS 1000.00 MB TO',F8.2,' MB =
1ETERS')
19 FORMAT ('0',//,T10,'NUMBER OF CASES 15N TC 25N',I5,8X,
1ASES 25N TO 35N',I5,8X,'NUMBER CF CASES 35N TO 45N',I5
20 FORMAT (2F6.1,7F9.2)
21 FORMAT (2F6.1,8F8.1)
C
            END
```

# SAMPLE OUTPUT FROM THICKNESS PROGRAM

60 LEVEL P	PROFILE	N. L	AT 19.2	W. LONG	63.0				
123456788911223144	RESS 1010 1015 1026 1026 1030 1040	TEMP 179-140 184-649 190-158 195-667 201-009 207-046 207-046 211-115 211-125 211-125 227-411 227-237-247	LEVEL 167 1189 1222 1234 1222 1234 1234 1234 1234 1234	PRESS 0.300 0.400 0.500 0.600 0.700 0.900 1.000 2.500 2.500 4.000 5.000 5.000	6693699012327 1227275899012327 1255345454568990123227 2555555555555555555555555555555555	E123456789012345	PRESSO 78.5000 90.00000 105.00000 105.00000 105.00000 105.00000 105.0000 105.0000 105.0000 105.0000 105.0000 105.0000 105.00000 105.0000 1	TEMP 240.953 237.450 236.283 233.9485 223.445 223.445 223.445 213.197 208.588 204.1706 199.241 198.009	LEVEL PRESS 100.000 47 150.000 48 250.000 51 400.000 52 450.000 53 500.000 554 550.000 556 650.000 557 700.000 559 900.000 60 1000.000
THICKNESS THICKNESS	150.00	MB TO	100.00 MB = 150.00 MB =	2398.78 13800.87					
THICKNESS THICKNESS	200.00	MB TO	100.00 MB = 200.00 MB =	4176.99 12022.65	METERS METERS				
THICKNESS	250.00 1000.00	MB TO	100.00 MB = 250.00 MB =	5620.17 10579.47					
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Radiance VTPR	mess mittance ple Regression	RMS errors					
An iterative technique is nesses of selected atmospheri radiance measurements. Layer atmospheric model are retriev radiative transfer equation, thicknesses of key atmospheri pressure levels. The retriev	developed for c layers from mean temperated by direct sand are then c layers bound	VTPR "clear-column" cures for a simplified solution of the ased to compute ded by commonly used					

#### (20. ABSTRACT continued)

reference wave numbers that vary from layer to layer. Transmittance tuning is employed to correct systematic errors in the retrieved mean temperatures. Thicknesses of key layers retrieved by the technique from "clear-column" radiances observed during a 24 hour period at scan spots between 15 N and 45 N are separated into three latitude-band samples. Each sample is subjected to stepwise multiple regression analysis to determine the thickness-specification of various standard layers in terms of the clear column radiances. RMS error-analyses resulting from the regression are then used to determine the quality of thickness-specifications of simulated tropospheres and stratospheres.

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